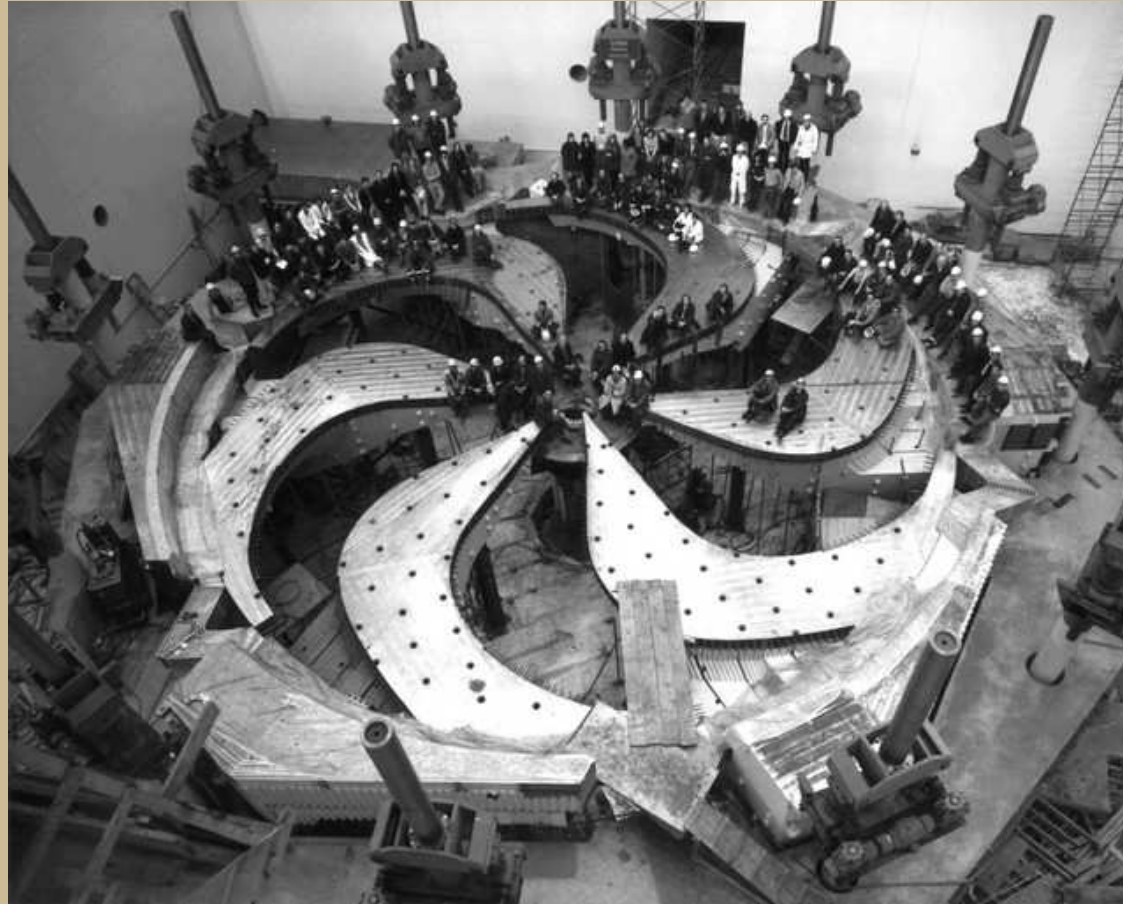


# Weak (and ridiculously weak) interactions as tests of fundamental physics



**Dan Melconian**

Cyclotron Institute/Dept. of Physics & Astronomy  
Texas A&M University

Presentation for REUs — July 25, 2012

# Overview

## 1. Fundamental symmetries

- what is our **current understanding**?
- what lies **beyond**?

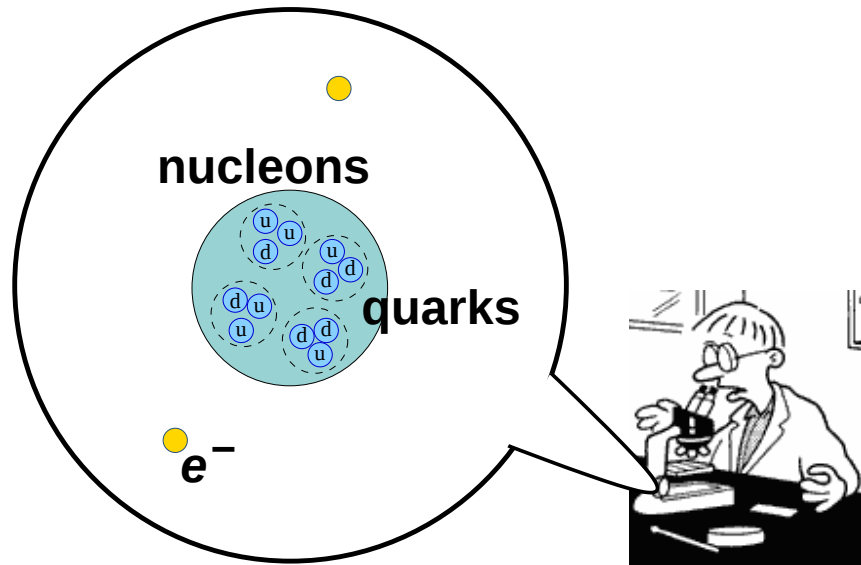
## 2. Tools of the trade

- **trapping** short-lived neutral atoms
- **polarizing** the atom cloud

## 3. Angular correlations using laser-cooled atoms

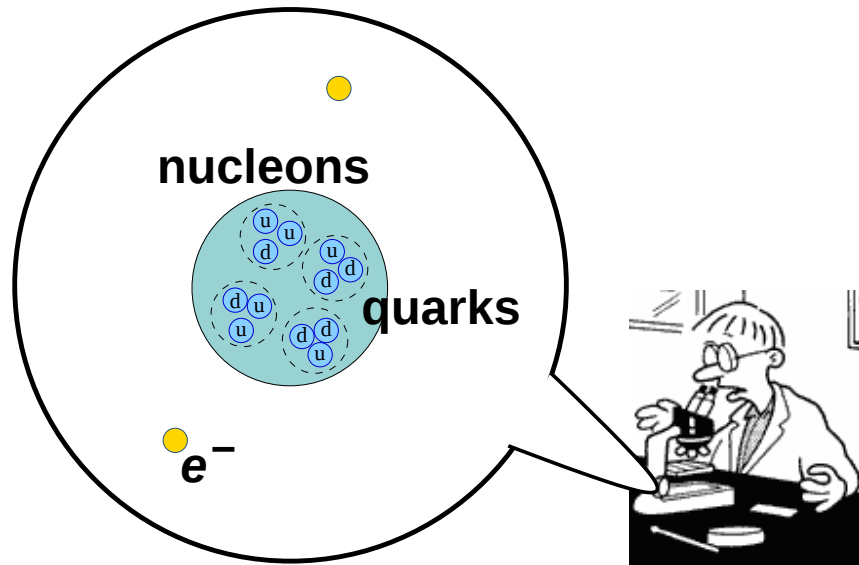
- angular correlations of **polarized  $^{37}\text{K}$**
- expected limits on right-handed currents

# Scope of fundamental physics

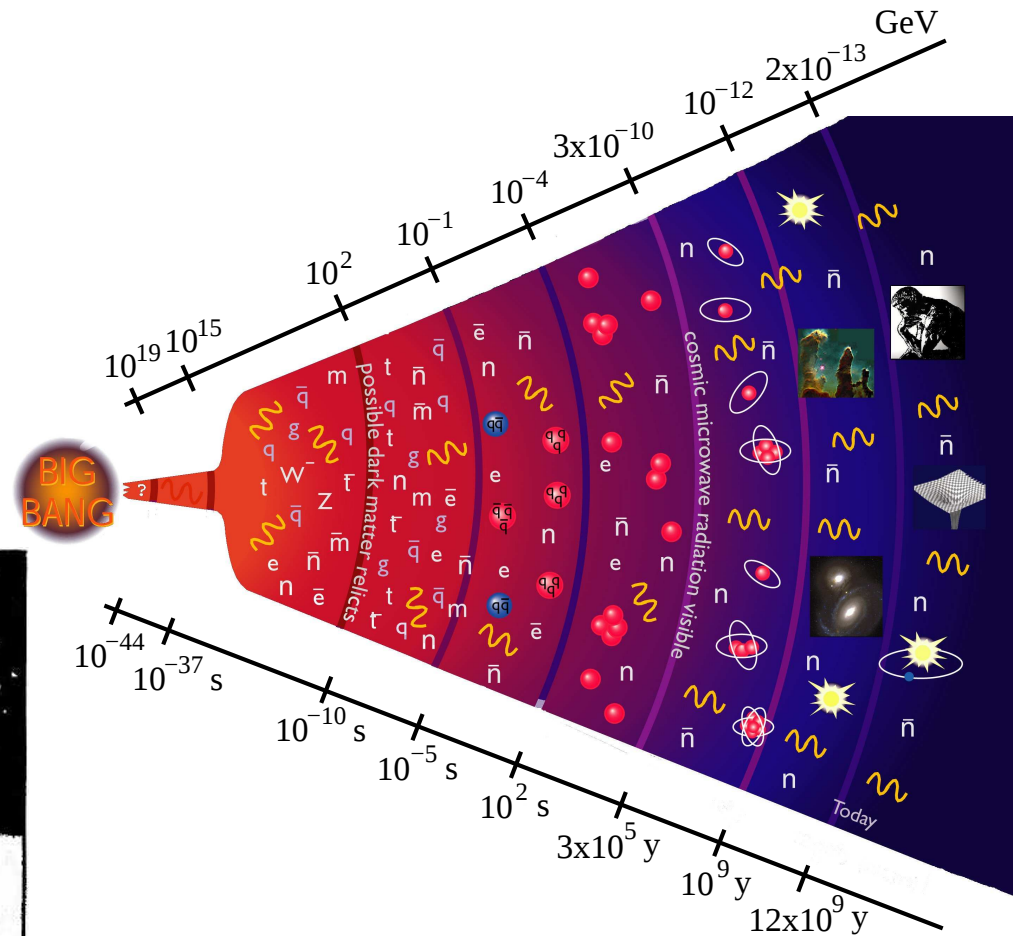
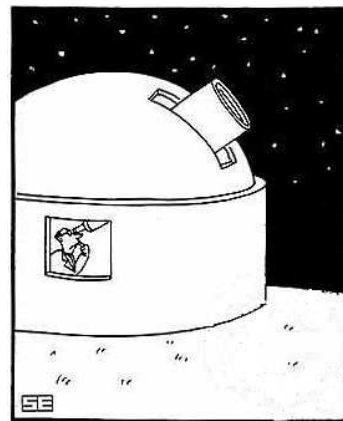


the atom  
from the very **smallest** scales ...

# Scope of fundamental physics



the atom  
from the very **smallest** scales ...



... to the very **largest**

# *The Standard Model*

All of the *known* elementary particles and their interactions are described within the framework of

**THE <sup>new</sup> STANDARD MODEL**

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and there's other symmetries too:

|           |                   |                  |
|-----------|-------------------|------------------|
| time      | $\Leftrightarrow$ | energy           |
| space     | $\Leftrightarrow$ | momentum         |
| rotations | $\Leftrightarrow$ | angular momentum |
| $\vdots$  |                   |                  |

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- $$\underbrace{SU(3)}_{\text{strong}} \times \overbrace{SU(2)_L \times U(1)}^{\text{electroweak}} + \underbrace{(\text{classical general rel})}_{\text{gravity}}$$

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## THE <sup>new</sup> STANDARD MODEL

- **quantum** + **special rel**  $\Rightarrow$  quantum field theory
- Noether's theorem: symmetry  $\Leftrightarrow$  conservation law
- $SU(3) \times SU(2)_L \times U(1)$ : strong + electroweak
- **12 elementary particles** and **4 fundamental forces**

|                | 1 <sup>st</sup>                            | 2 <sup>nd</sup>                                | 3 <sup>rd</sup>                                  | $Q$  | mediator | force  |
|----------------|--------------------------------------------|------------------------------------------------|--------------------------------------------------|------|----------|--------|
| <b>leptons</b> | $\begin{pmatrix} \nu_e \\ e \end{pmatrix}$ | $\begin{pmatrix} \nu_\mu \\ \mu \end{pmatrix}$ | $\begin{pmatrix} \nu_\tau \\ \tau \end{pmatrix}$ | 0    | $g$      | strong |
|                |                                            |                                                |                                                  | -1   | $W^\pm$  | weak   |
| <b>quarks</b>  | $\begin{pmatrix} u \\ d \end{pmatrix}$     | $\begin{pmatrix} c \\ s \end{pmatrix}$         | $\begin{pmatrix} t \\ b \end{pmatrix}$           | +2/3 | $Z^0$    |        |
|                |                                            |                                                |                                                  | -1/3 | $\gamma$ | EM     |

*That's all fine and dandy, but...*

Does the Standard Model work??

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- **predicted** the existence of the  $W^{\pm}$ ,  $Z_0$ ,  $g$ ,  $c$  and  $t$

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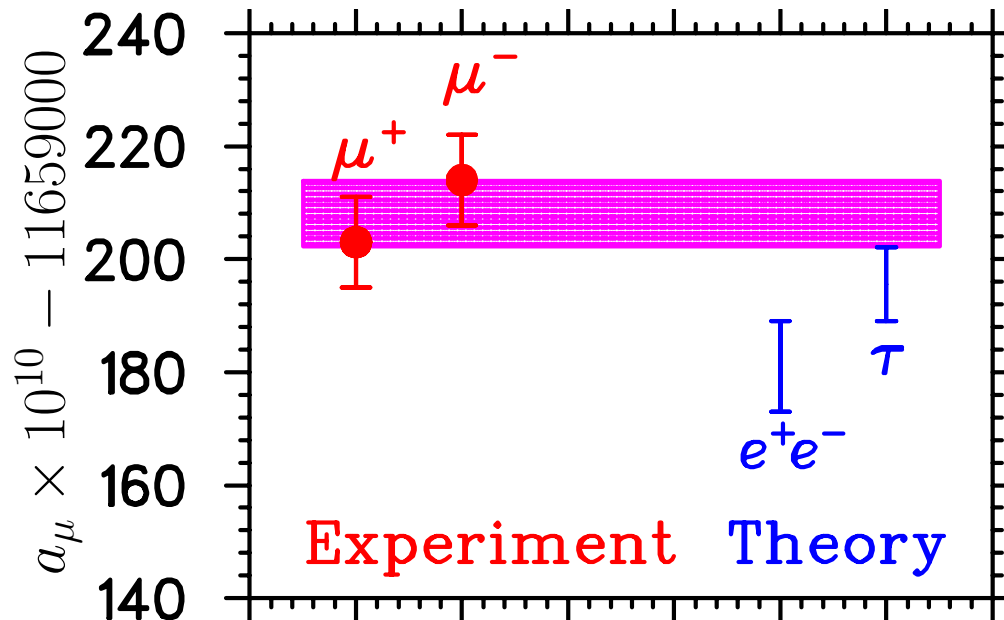
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$$a_\mu \equiv \frac{1}{2}(g - 2)$$

$$a_\mu(\text{exp}) = 11\,659\,208(6) \times 10^{-10}$$

$$a_\mu(\text{SM}) = 11\,659\,181(8) \times 10^{-10}$$

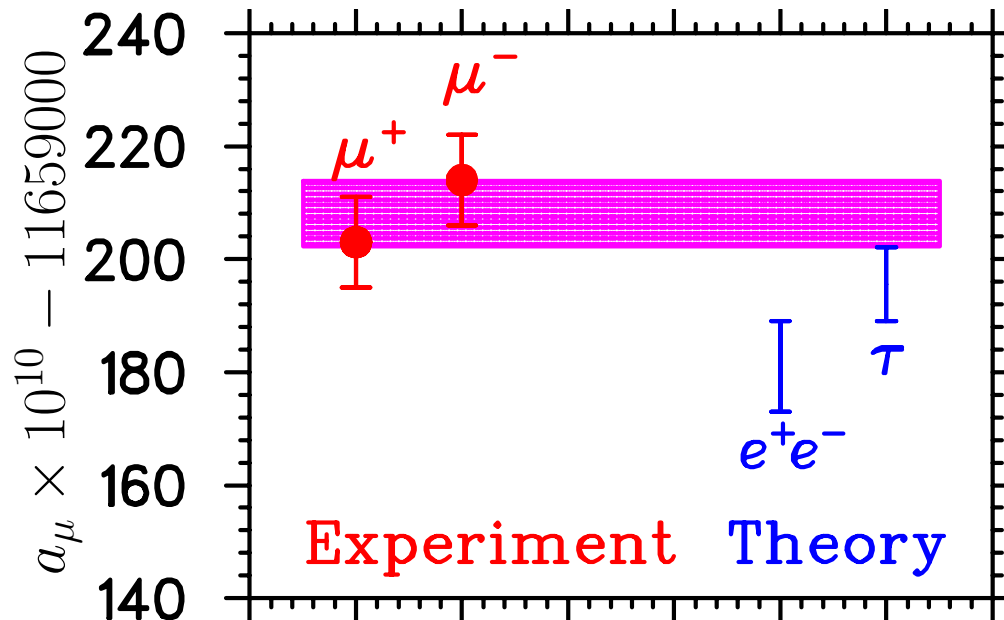
**$\pm 1$  part-per-million!!**

(PRL **92** (2004) 161802)

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Wow ... this is

the most precisely tested theory ever conceived!

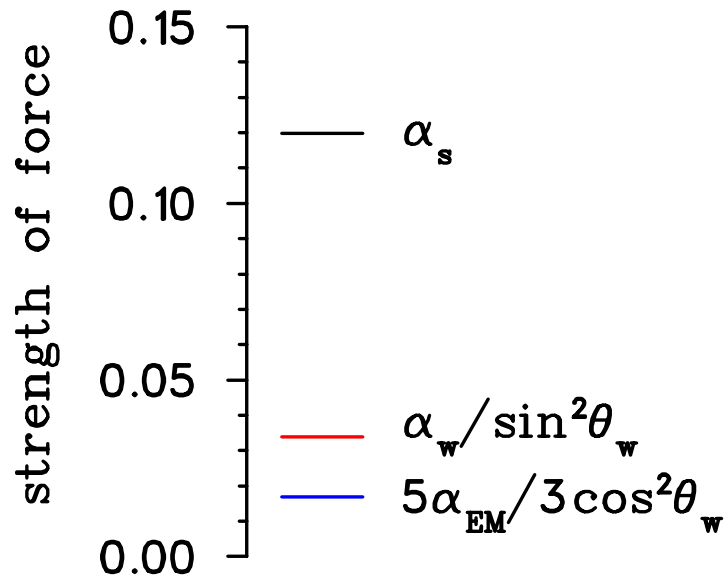


# *But there are still questions ...*

- values of parameters: does our “ultimate” theory *really* need **25** arbitrary constants? Do they **change** with time?
- dark matter: SM physics makes up **only 4%** of the energy-matter of the universe!
- baryon asymmetry: why more **matter** than **anti-matter**?
- strong  $CP$ : do **axions** exist? **Fine-tuning**?
- neutrinos: **Dirac** or **Majorana**?
- fermion generations: why **three** families?
- weak mixing: Is the CKM matrix **unitary**?
- parity violation: is parity **maximally** violated in the weak interaction? No **right-handed** currents?
- EW symmetry breaking: how do the fermions acquire **mass**? Mass **hierarchy**?
- gravity: of course can't forget about a **quantum** description of **gravity**!

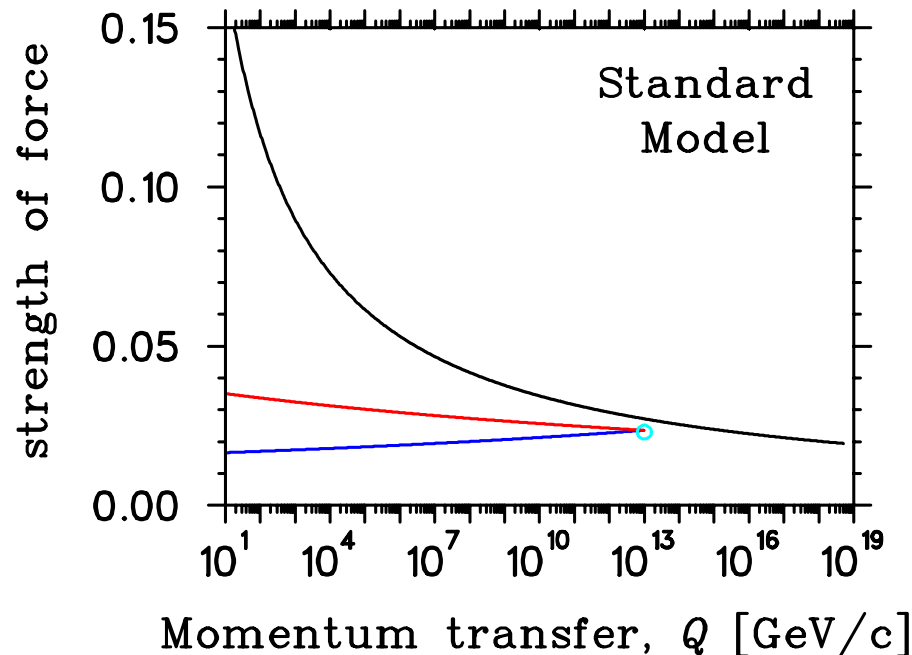
# Beyond the Standard Model

At our energy scales, we see four distinct forces ...



# Beyond the Standard Model

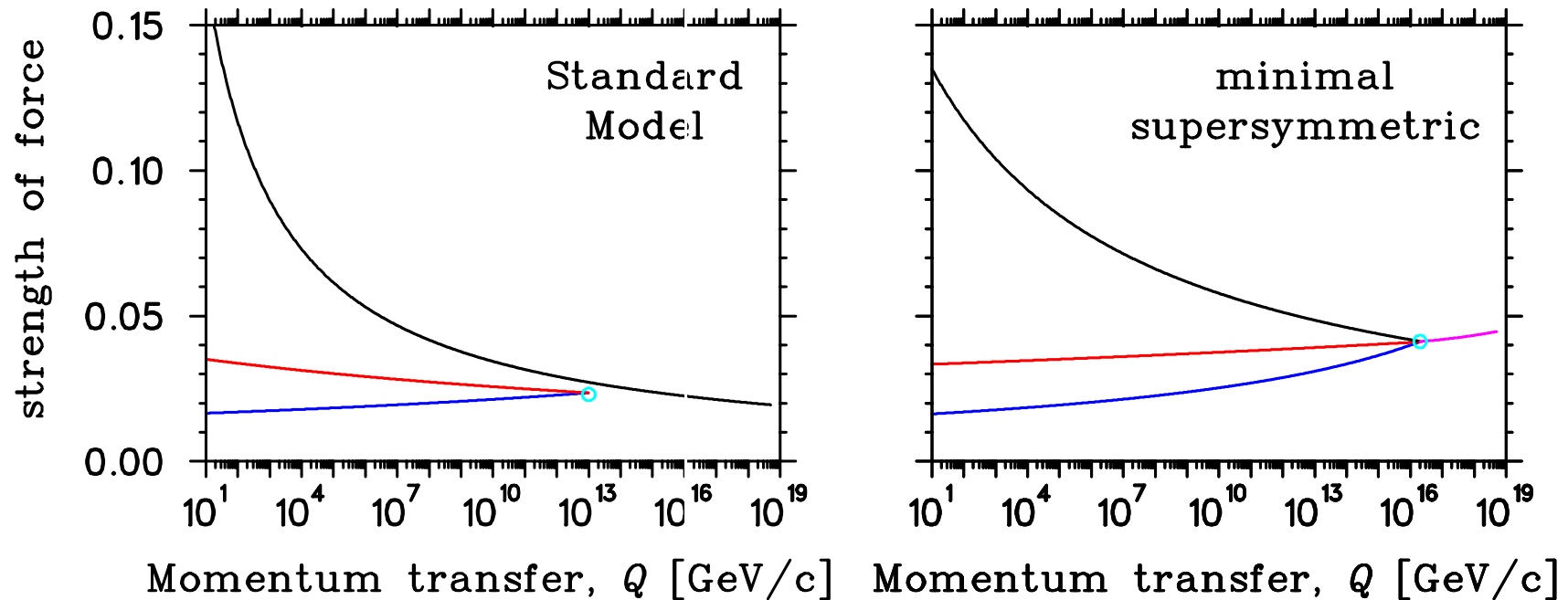
But these coupling ‘constants’ **aren’t** really constant:  $\alpha_i \rightarrow \alpha_i(Q)$



- electromagnetic and weak strengths equal at  $\approx 10^{13}$  GeV
- strong force gets weaker, but doesn't unify with EW....

# Beyond the Standard Model

But what if there is **new physics** we haven't seen yet?



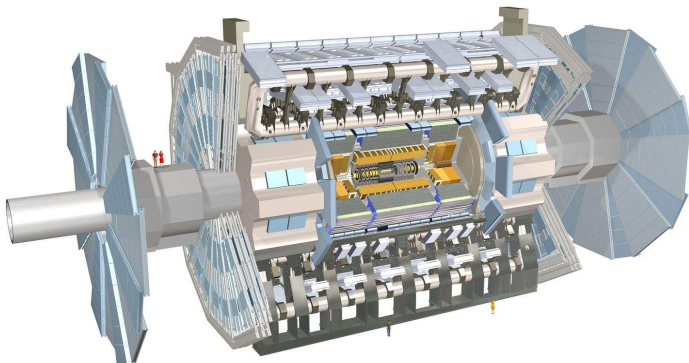
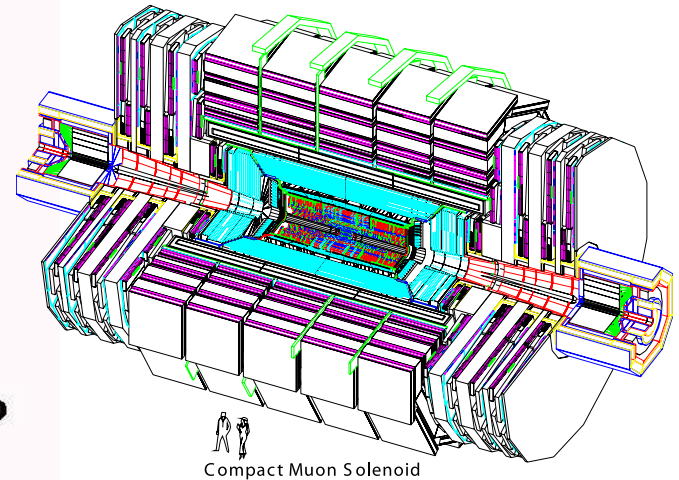
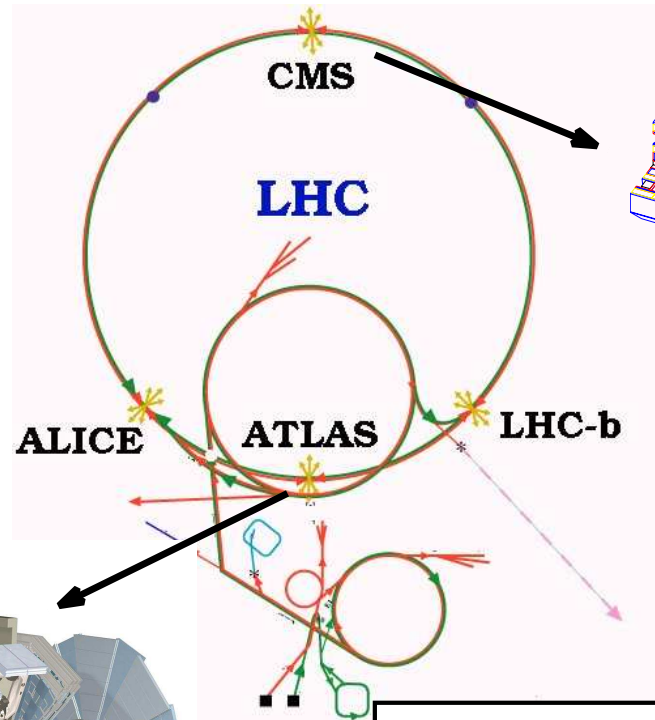
the running of the coupling constants would be affected;  
maybe they converge at some GUT scale?

Are the three theories of **E & M**, **weak** and **strong**  
interactions all **low-energy limits** of  
**one unifying** theory?

# How do we test the SM?

**colliders:** CERN, SLAC, FNAL, BNL, KEK, DESY, ....

**direct** search of particles



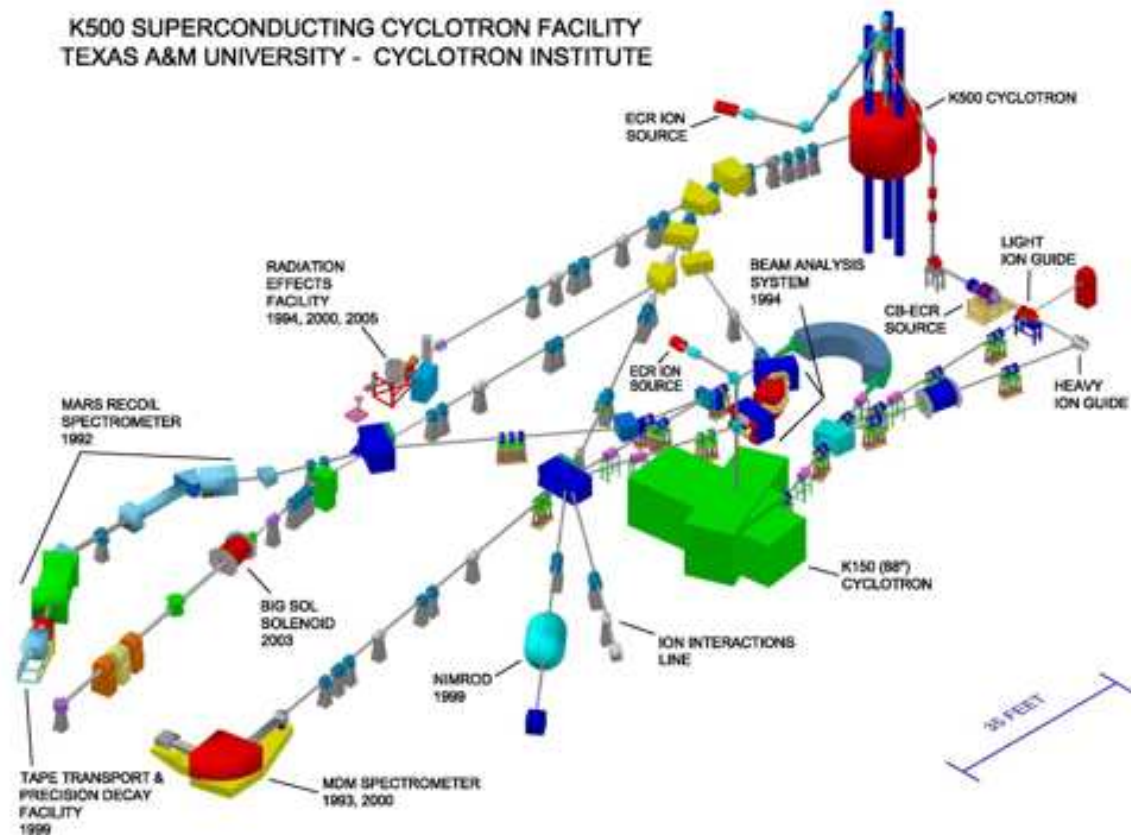
***“go big or go home”***

- large multi-national collabs
- billion \$ price-tags 🤪

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**nuclear physics:** radioactive ion beam facilities (ISOL/frag)

**indirect** search via precision measurements

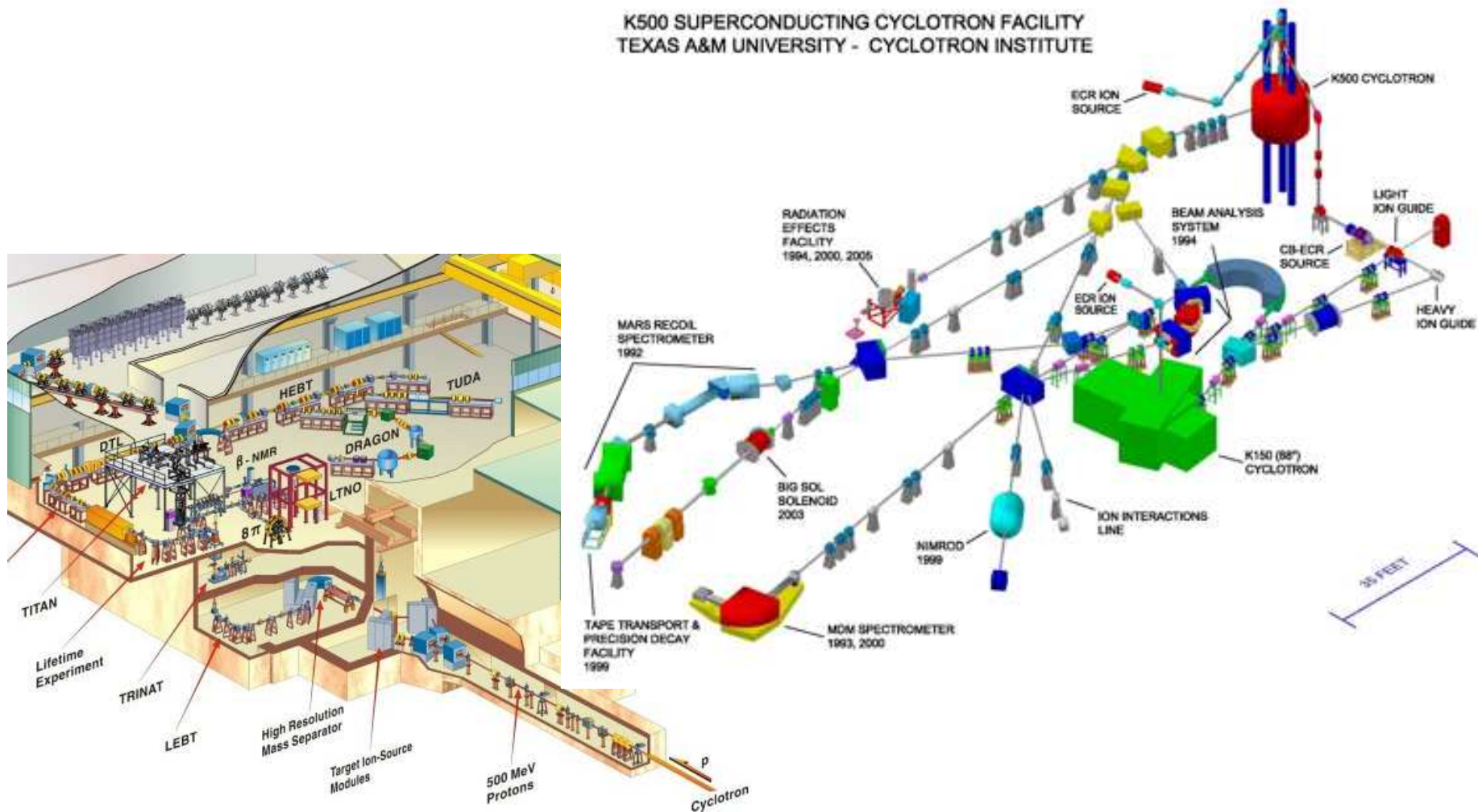


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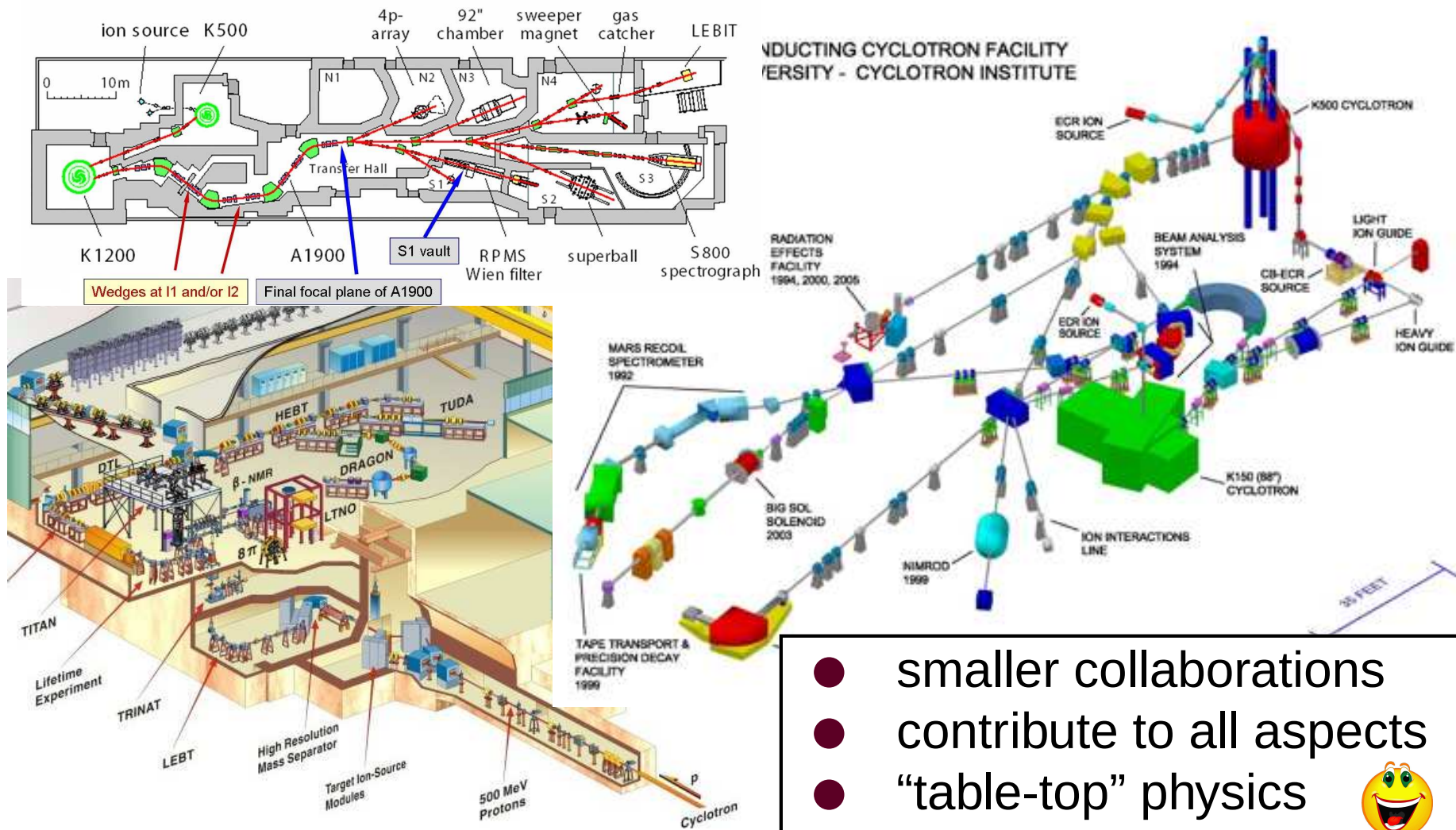
K500 SUPERCONDUCTING CYCLOTRON FACILITY  
TEXAS A&M UNIVERSITY - CYCLOTRON INSTITUTE



# How do we test the SM?

nuclear physics: radioactive ion beam facilities (ISOL/frag)

**indirect** search via precision measurements



- smaller collaborations
- contribute to all aspects
- “table-top” physics 😄

# How we all test the SM

- **colliders**: CERN, SLAC, FNAL, BNL, KEK, DESY ...
- **nuclear physics**: traps, exotic beams, neutron, EDMs,  $0\nu\beta\beta$ , ...
- **cosmology & astrophysics**: SN1987a, Big Bang nucleosynthesis, ...
- **muon decay**: Michel parameters:  $\rho$ ,  $\delta$ ,  $\eta$ , and  $\xi$
- **atomic physics**: anapole moment, spectroscopy, ...

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all of these techniques are **complementary** and **important**

- different experiments probe different (new) physics
- if signal seen, cross-checks crucial!

often they are **interdisciplinary**

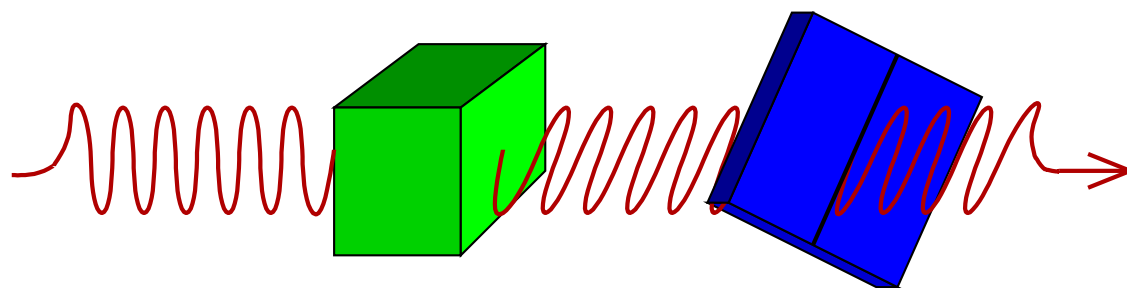
(fun and a great basis for graduate students!)

# *History up to 1957*

- 1768: Kant debates nature of incongruent counterparts

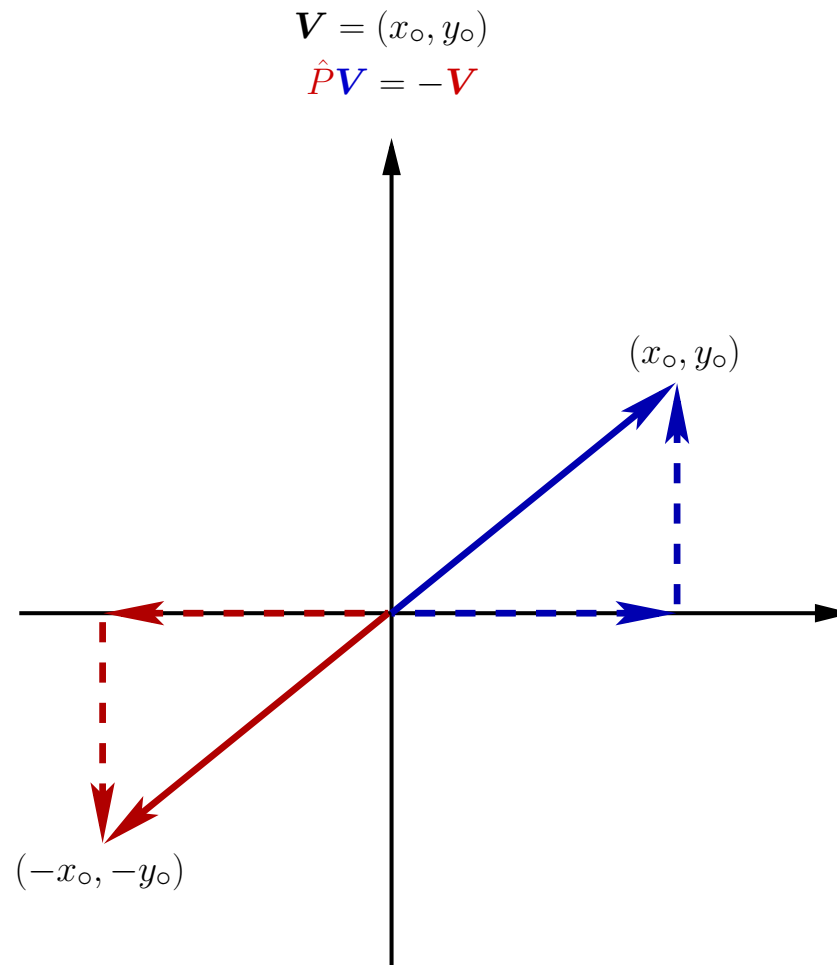
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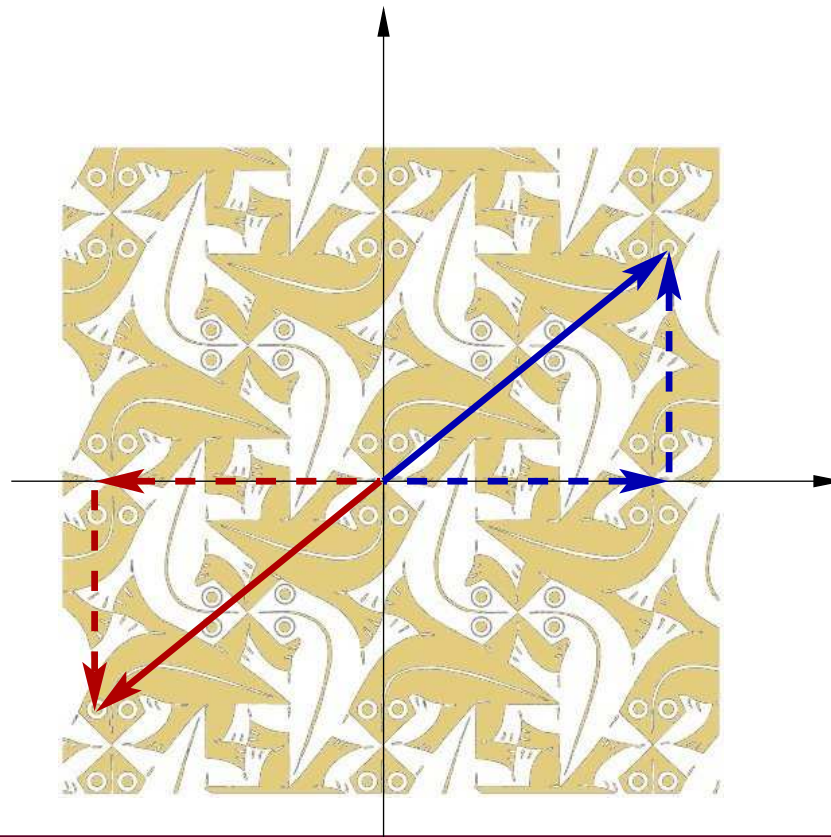


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$A$  = M.C. Escher reptiles

$$\hat{P}A = +A$$



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intrinsic parity  $\Leftrightarrow$  helicity or “handedness”



$$\mathbf{r} \times \mathbf{p} \rightarrow (-\mathbf{r}) \times (-\mathbf{p}) = \mathbf{r} \times \mathbf{p}$$

(spin is an **axial** vector)

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“Left-handed”



“Right-handed”

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- 1934: Fermi’s theory of  $\beta$  decay:

$$\frac{dW}{dE_e} = \frac{G_F^2}{(2\pi)^5} p_e E_e (E_e - A_o)^2 |\mathcal{M}_{fi}|^2$$

$$\mathcal{M}_{fi} = \int \psi^* \Gamma_i \psi, \quad \Gamma_i = (S, P, V, A, T)$$

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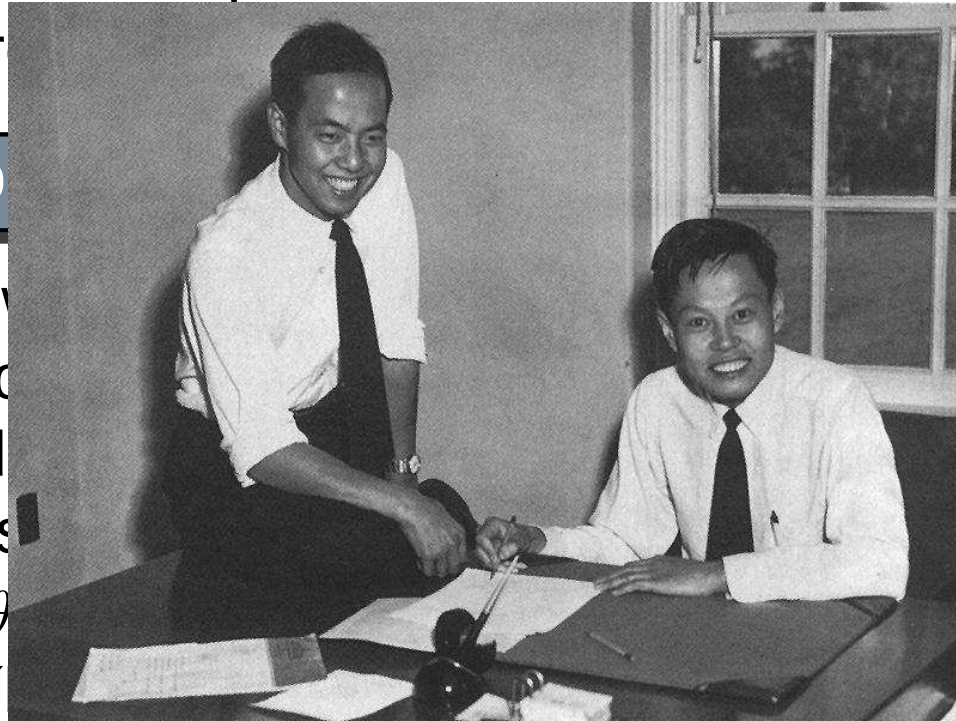
$\theta^+ \rightarrow \pi^0 \pi^+$  and  $\tau^+ \rightarrow \pi^+ \pi^+ \pi^-$   
(but **same** lifetime, mass, ...)

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intrinsic p

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- 1953:  $^6\text{He}$ ,  $^{19}\text{Ne}$  decays **(S, T)**
- 1953: Dalitz points out  $\theta$  (
- 1956: prompted Lee and Yang to **question** current convention



... existing experiments do indicate **parity conservation in strong and electromagnetic interactions**, but that for **weak interactions** ... parity conservation is so far only an **extrapolated hypothesis** unsupported by experimental evidence.

(Feynman bets parity is conserved)

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$$K^+ \rightarrow \pi^0 \pi^+ \text{ and } K^+ \rightarrow \pi^+ \pi^+ \pi^-$$

(**same** particle; parity **not** conserved)

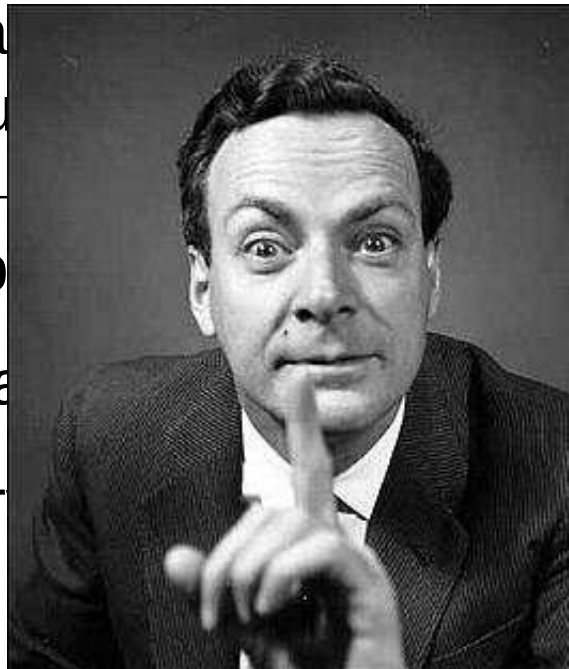
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- 1956: prompted Lee and Yang to question current convention
- 1957: Wu ( ${}^{60}\text{Co}$ ), Garwin ( ${}^{60}\text{Co} \rightarrow e^+ \nu_e \gamma$ )  $- A$ ) interaction



(Feynman loses \$50)

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 $K^+ \rightarrow \pi^0 e^+ \nu_e$  (S, T)  $\theta - \tau$  puzzle  
 $\beta$  decay has a long history of developing our understanding of fundamental symmetries
- 1956: Lee and Yang to **question** current convention (P not conserved)
- 1957: Wu ( ${}^{60}\text{Co}$ ), Garwin ( $\mu^+$ ) favour **(V – A) interaction**

# The Standard Model (and beyond)

The **Electroweak Interaction**:  $SU(2)_L \times U(1) \Rightarrow W_L^\pm, Z^0, \gamma$

Built upon **maximal** parity violation:

Vector  $\hat{P}|\Psi\rangle = -|\Psi\rangle$

$$H_{\text{SM}} = G_F V_{ud} \bar{e} (\gamma_\mu - \gamma_\mu \gamma_5) \nu_e \bar{u} (\gamma^\mu - \gamma^\mu \gamma_5) d$$

Axial – vector  $\hat{P}|\Psi\rangle = +|\Psi\rangle$

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low-energy limit of a **deeper**  $SU(2)_R \times SU(2)_L \times U(1)$  theory?

$\Rightarrow$  3 more vector bosons:  $W_R^\pm, Z'$

Simplest extensions: “*manifest left-right symmetric*” models

$\rightsquigarrow$  only 2 new parameters:  $W_2$  mass and a mixing angle,  $\zeta$

$$|W_L\rangle = \cos \zeta |W_1\rangle - \sin \zeta |W_2\rangle$$

$$|W_R\rangle = \sin \zeta |W_1\rangle + \cos \zeta |W_2\rangle$$

# *RHCs would affect correlation parameters*

$$A_\beta = \frac{-2\rho}{1+\rho^2} \left( \sqrt{\frac{3}{5}} - \frac{\rho}{5} \right)$$

$$B_\nu = \frac{-2\rho}{1+\rho^2} \left( \sqrt{\frac{3}{5}} + \frac{\rho}{5} \right)$$

and  $R_{\text{slow}} = 0$

# RHCs would affect correlation parameters

In the presence of **new physics**, the **angular distribution of  $\beta$  decay** will be affected.

$$A_\beta = \frac{-2\rho}{1+\rho^2} \left( \sqrt{\frac{3}{5}} - \frac{\rho}{5} \right) \rightarrow \frac{-2\rho}{1+\rho^2} \left[ (1-xy) \sqrt{\frac{3(1+x^2)}{5(1+y^2)}} - \frac{\rho(1-y^2)}{5(1+y^2)} \right]$$

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and  $R_{\text{slow}} = 0 \rightarrow y^2$

where  $x \approx (M_L/M_R)^2 - \zeta$  and  $y \approx (M_L/M_R)^2 + \zeta$

are RHC parameters that are zero in the SM.

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$$B_\nu = \frac{-2\rho}{1+\rho^2} \left( \sqrt{\frac{3}{5}} + \frac{\rho}{5} \right) \rightarrow \frac{-2\rho}{1+\rho^2} \left[ (1-xy) \sqrt{\frac{3(1+x^2)}{5(1+y^2)}} + \frac{\rho(1-y^2)}{5(1+y^2)} \right]$$

and  $R_{\text{slow}} = 0 \rightarrow y^2$

where  $x \approx (M_L/M_R)^2 - \zeta$  and  $y \approx (M_L/M_R)^2 + \zeta$

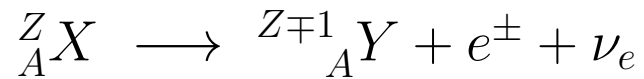
are RHC parameters that are zero in the SM.

$\Rightarrow$  Precision measurements test the SM

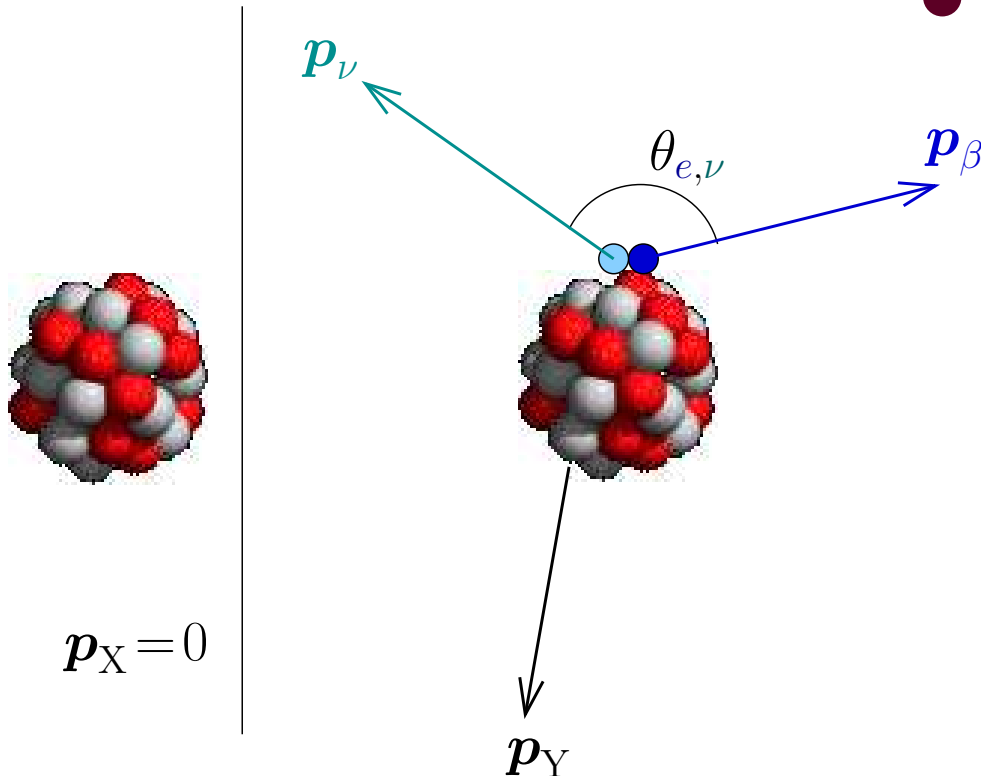
Goal must be  $\lesssim 0.1\%$

(see Profumo, Ramsey-Musolf and Tulin, PRD **75** (2007))

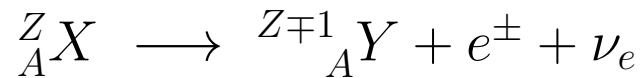
# The Gameplan:



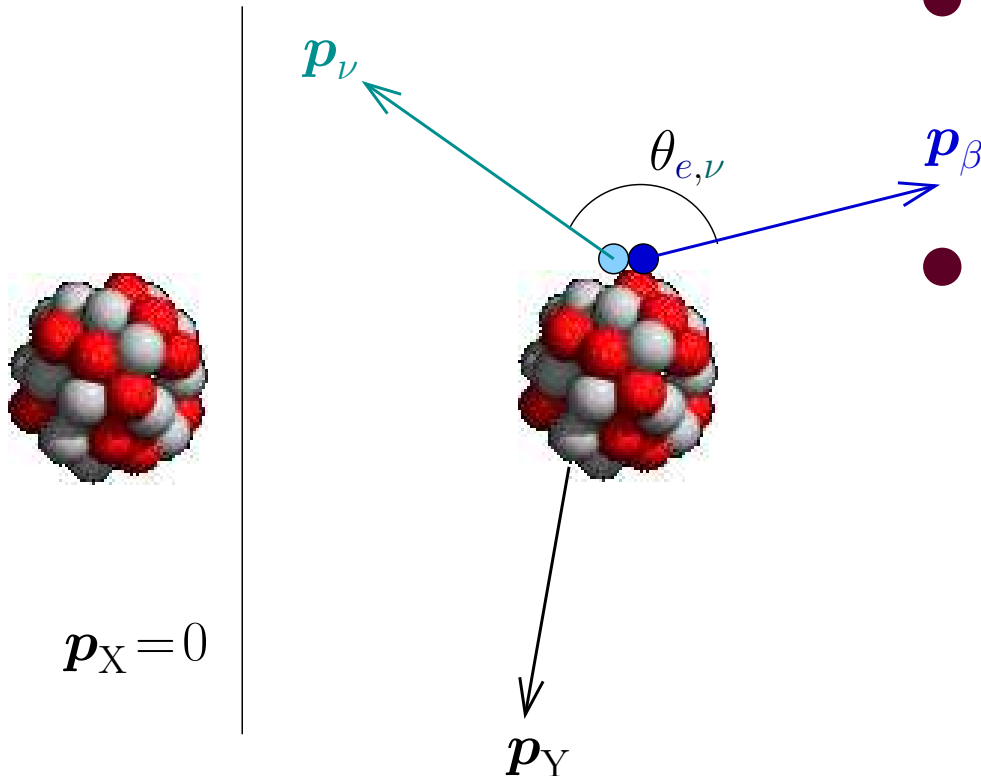
- perform a  $\beta$  decay experiment on **short-lived** isotopes



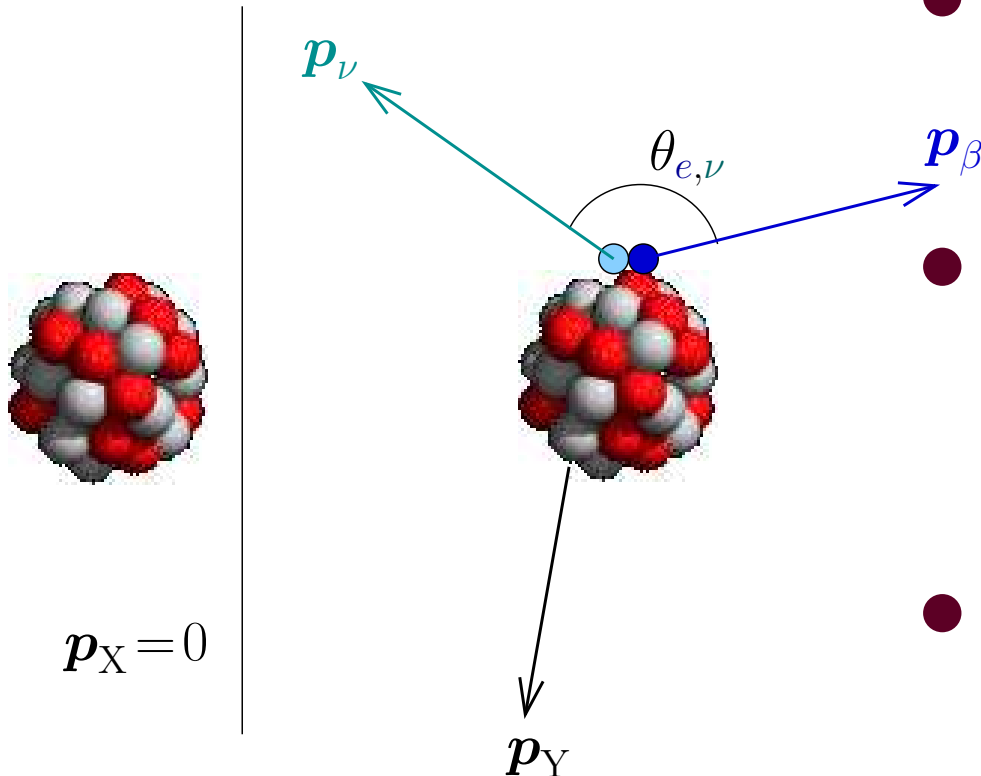
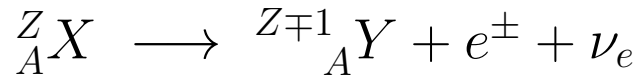
# The Gameplan:



- perform a  $\beta$  decay experiment on **short-lived** isotopes
- make a **precision measurement** of the angular correlation parameters

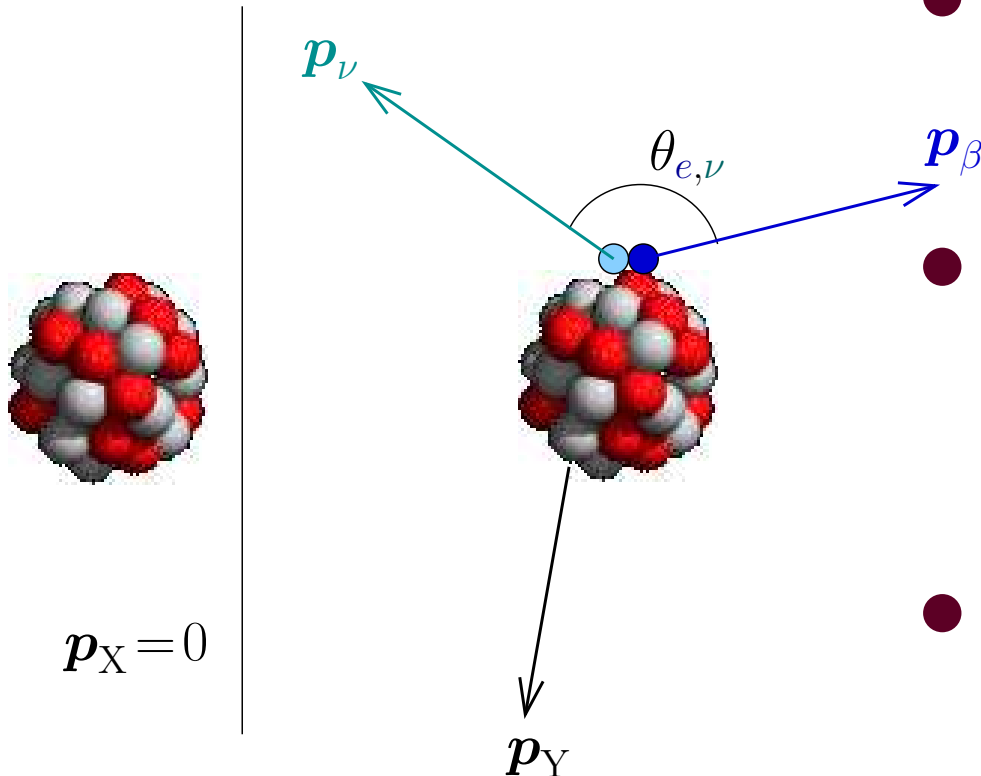
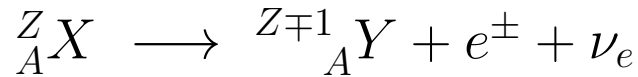


# The Gameplan:



- perform a  $\beta$  decay experiment on **short-lived** isotopes
- make a **precision measurement** of the angular correlation parameters
- **compare** the SM predictions to observations

# The Gameplan:



- perform a  $\beta$  decay experiment on **short-lived** isotopes
- make a **precision measurement** of the angular correlation parameters
- **compare** the SM predictions to observations
- look for **deviations** as an indication of **new physics**

# Wu's experiment

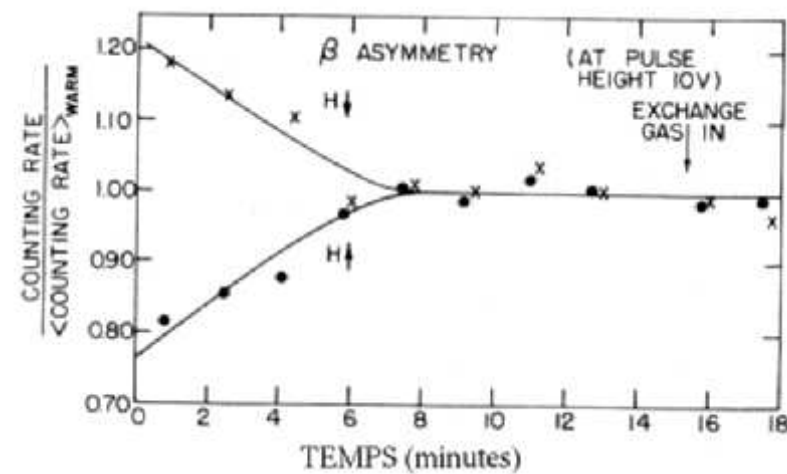
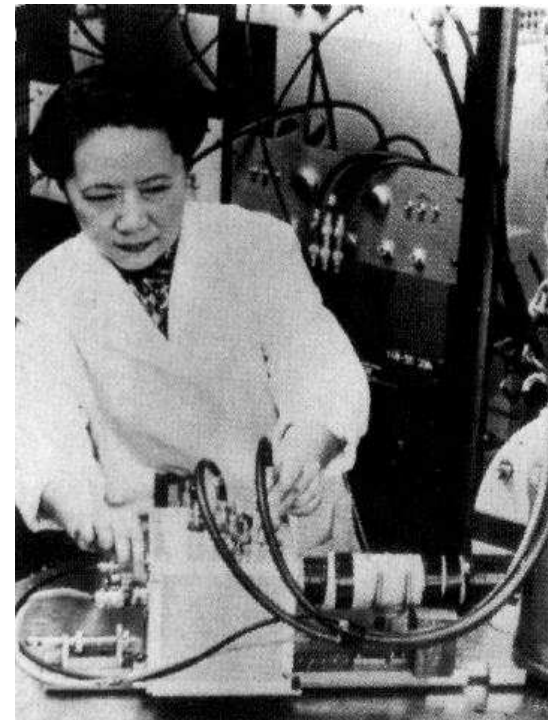
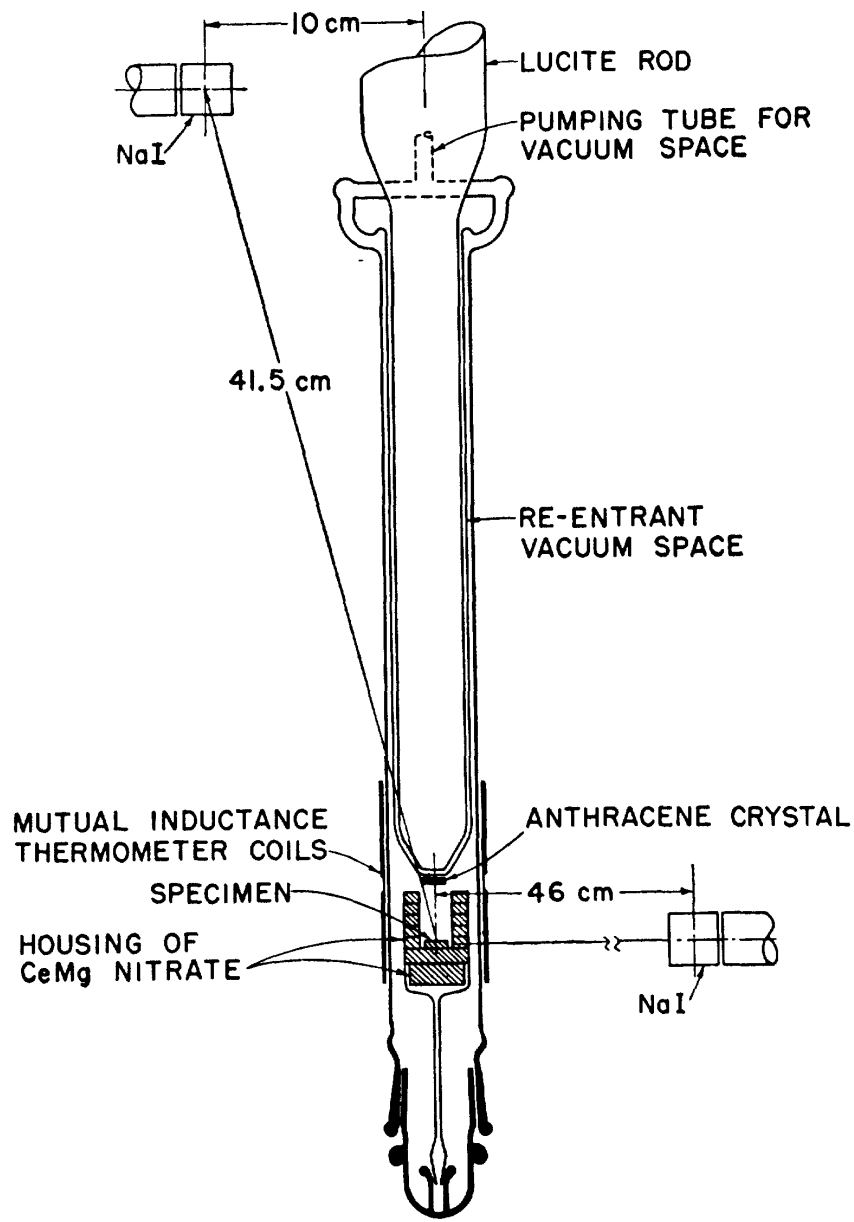
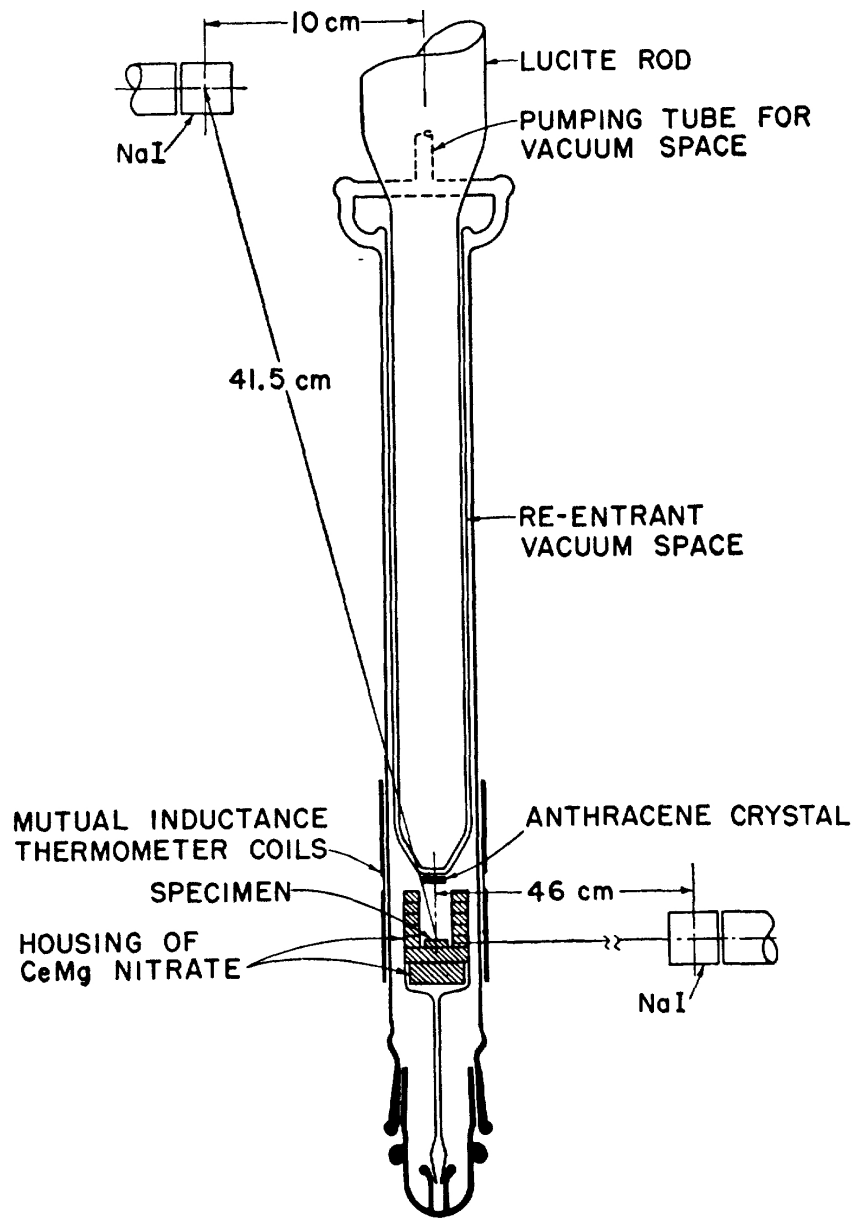


FIG. 1. Schematic drawing of the lower part of the cryostat.

# Wu's experiment



- so much scattering!
- low polarization
- short relaxation time
- poor sample purity
- pain to flip spin
- need long  $t_{1/2}$

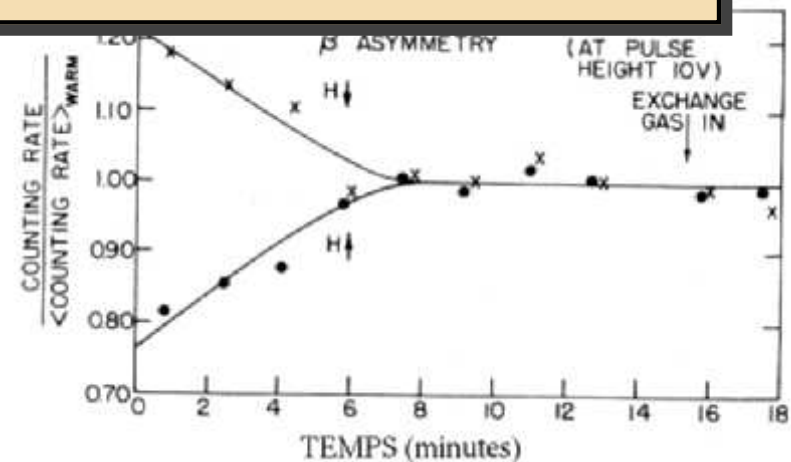
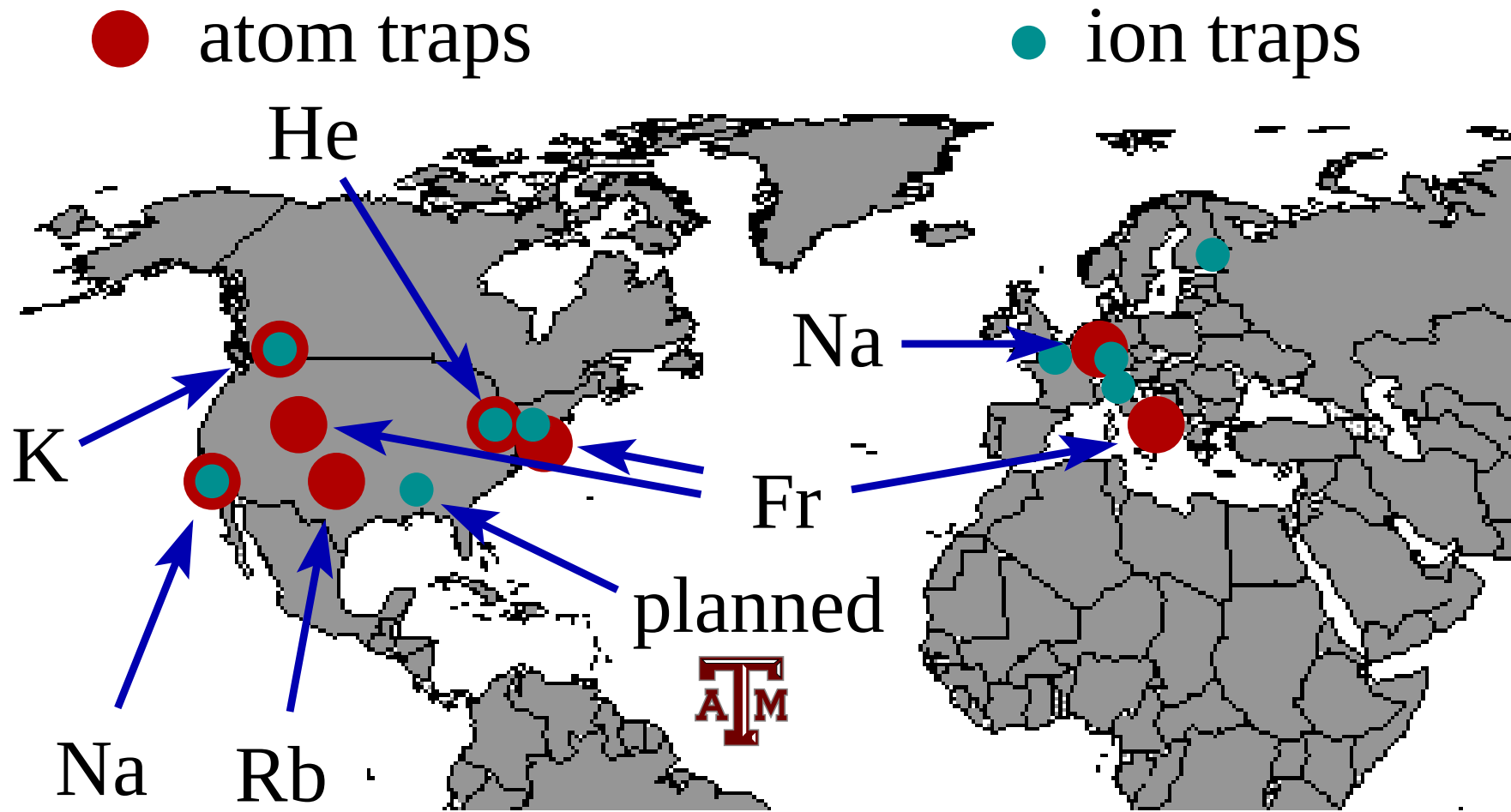


FIG. 1. Schematic drawing of the lower part of the cryostat.

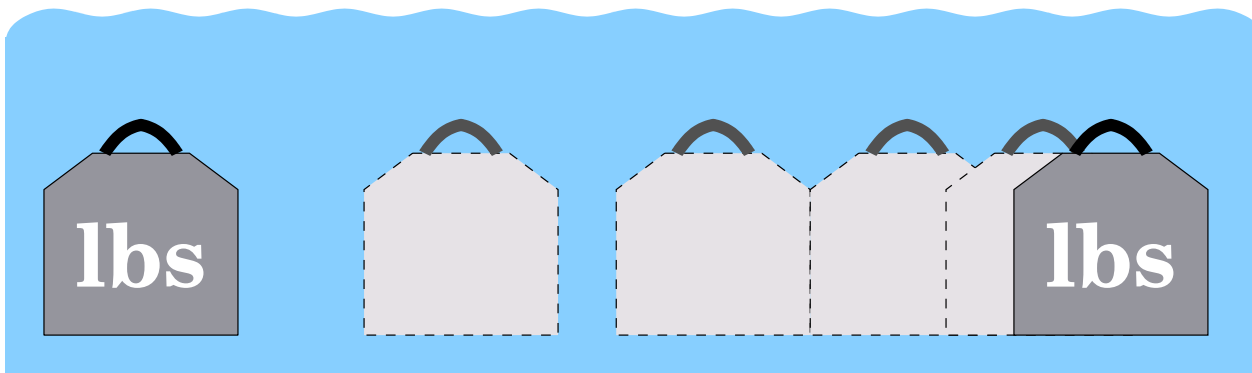
# Traps around the world

Many groups around the world realize the potential of using traps for precision weak interaction studies



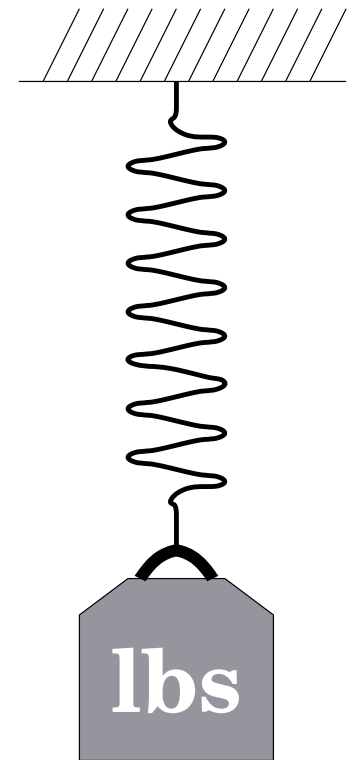
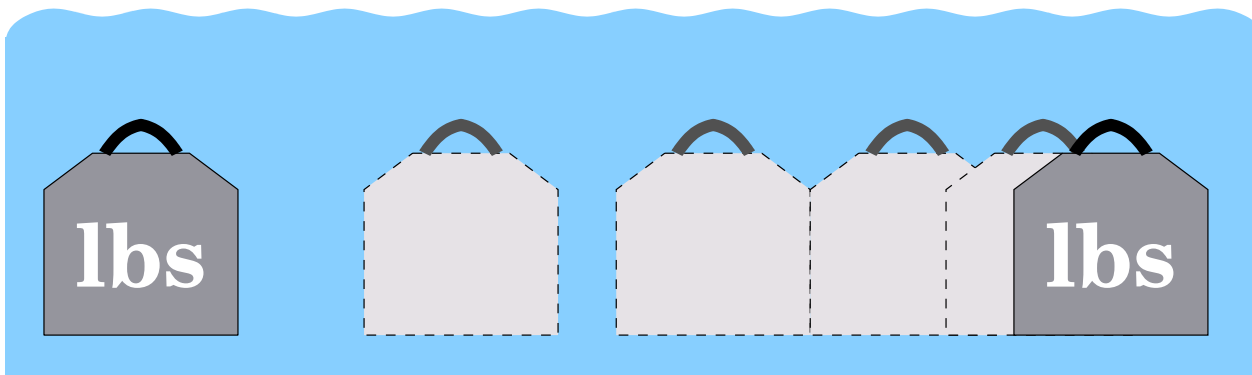
# Neutral Atom Traps

**Any** type of trap requires a **velocity-dependent** force to cool an object



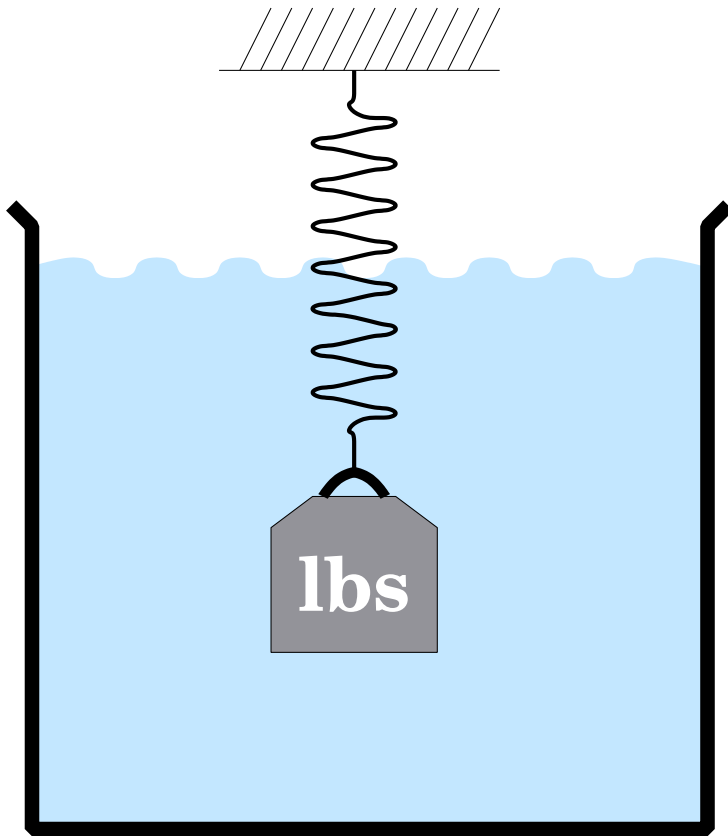
# Neutral Atom Traps

**Any** type of trap requires a **velocity-dependent** force to cool an object . . . as well as a **position-dependent** force that defines  $x = 0$



# Neutral Atom Traps

**Any** type of trap requires a **velocity-dependent** force to cool an object ... as well as a **position-dependent** force that defines  $x = 0$



Laser light

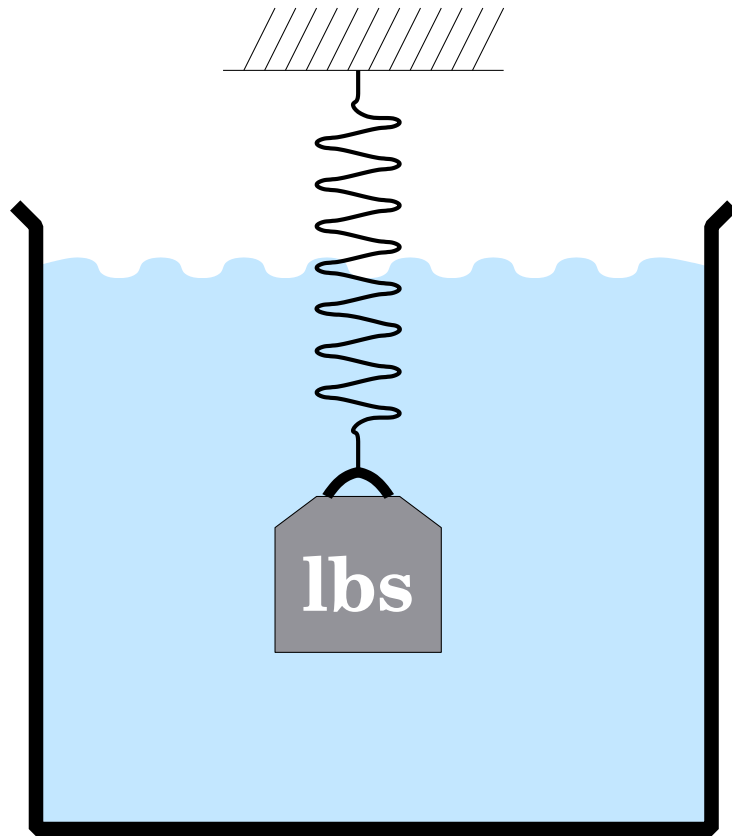
→ velocity-dependent force

Zeeman effect

→ position-dependent force

# Neutral Atom Traps

**Any** type of trap requires a **velocity-dependent** force to cool an object ... as well as a **position-dependent** force that defines  $x = 0$

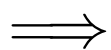


Laser light

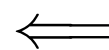
→ velocity-dependent force

Zeeman effect

→ position-dependent force



atom trap = damped harmonic oscillator

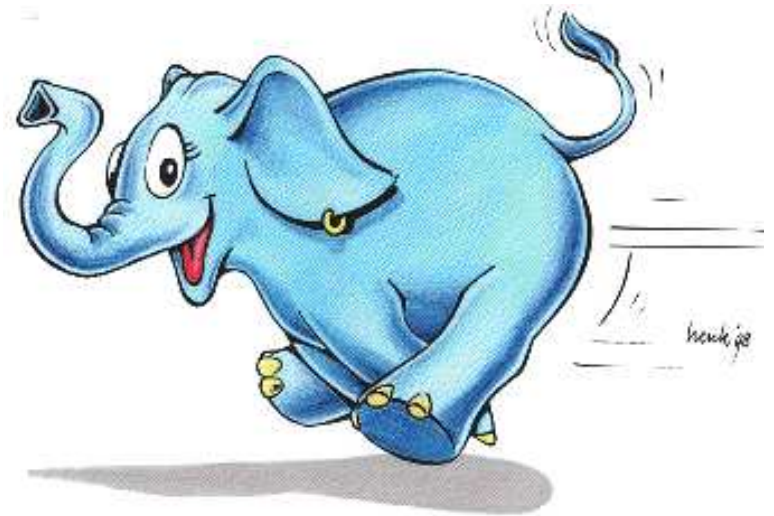


# Atom-Photon Interactions

How can light seriously affect a thermal atom?



VS.



$$\begin{aligned}\hbar c \cdot \frac{2\pi}{\lambda} &= (197.3 \text{ MeV fm}) \left( \frac{6.28}{770 \text{ nm}} \right) \\ &= 1.6 \times 10^{-6} \text{ MeV}\end{aligned}$$

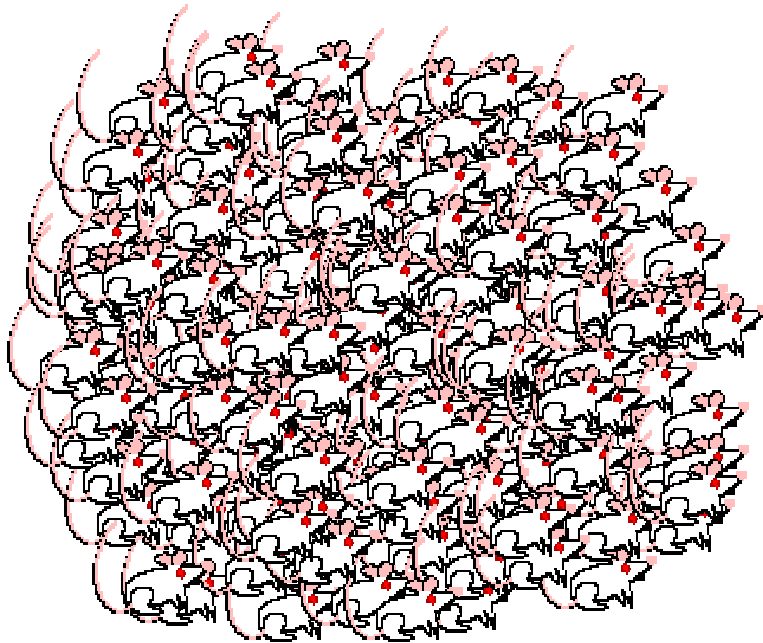
$$\Rightarrow \boxed{\hbar k \sim 1.6 \text{ eV/c}}$$

$$\begin{aligned}\frac{1}{2} M v^2 &= k_B T \\ M v &= \left[ 2(40 \times 10^6 \text{ keV/c}^2) \right. \\ &\quad \left. (8.62 \times 10^{-8} \text{ keV/K})(295 \text{ K}) \right]^{1/2}\end{aligned}$$

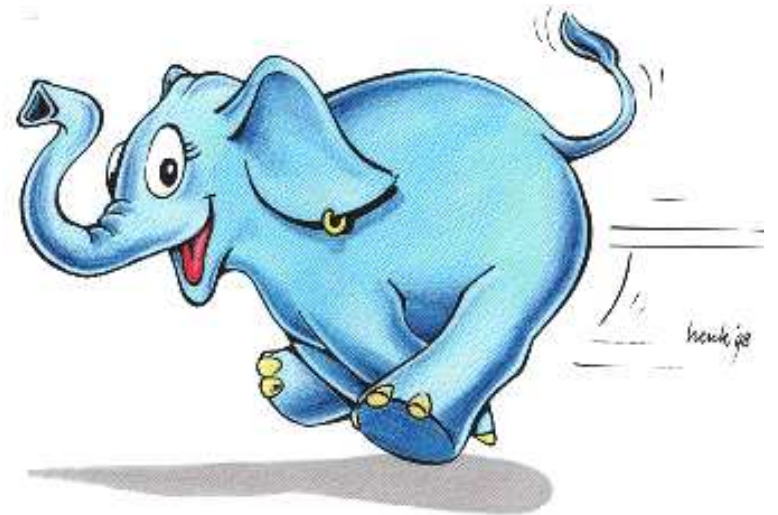
$$\Rightarrow \boxed{M v \sim 45 \text{ keV/c}}$$

# Atom-Photon Interactions

## *Cycling Transitions!*



VS.



$$\hbar k \times 30,000 \approx Mv$$

# However....

cycling transition  $\Rightarrow$  not everything trappable

|    |    |    |    |    |    |    |    |    |    |    |    |    |    |    |    |    |    |
|----|----|----|----|----|----|----|----|----|----|----|----|----|----|----|----|----|----|
| H  |    |    |    |    |    |    |    |    |    |    |    |    |    |    |    |    | He |
| Li | Be |    |    |    |    |    |    |    |    |    |    | B  | C  | N  | O  | F  | Ne |
| Na | Mg |    |    |    |    |    |    |    |    |    |    | Al | Si | P  | S  | Cl | Ar |
| K  | Ca | Sc | Ti | V  | Cr | Mn | Fe | Co | Ni | Cu | Zn | Ga | Ge | As | Se | Br | Kr |
| Rb | Sr | Y  | Zr | Nb | Mo | Tc | Ru | Rh | Pd | Ag | Cd | In | Sn | Sb | Te | I  | Xe |
| Cs | Ba | Lu | Hf | Ta | W  | Re | Os | Ir | Pt | Au | Hg | Tl | Pb | Bi | Po | At | Rn |
| Fr | Ra | Lr | Rf | Db | Sg | Bh | Hs | Mt | Ds | Rg |    |    |    |    |    |    |    |

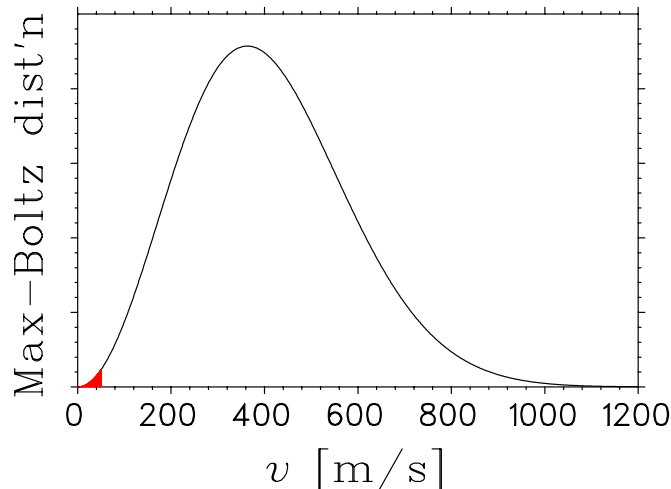
|    |    |    |    |    |    |    |    |    |    |    |    |    |    |
|----|----|----|----|----|----|----|----|----|----|----|----|----|----|
| La | Ce | Pr | Nd | Pm | Sm | Eu | Gd | Tb | Dy | Ho | Er | Tm | Yb |
| Ac | Th | Pa | U  | Np | Pu | Am | Cm | Bk | Cf | Es | Fm | Hd | No |

# However....

cycling transition  $\Rightarrow$  not everything trappable

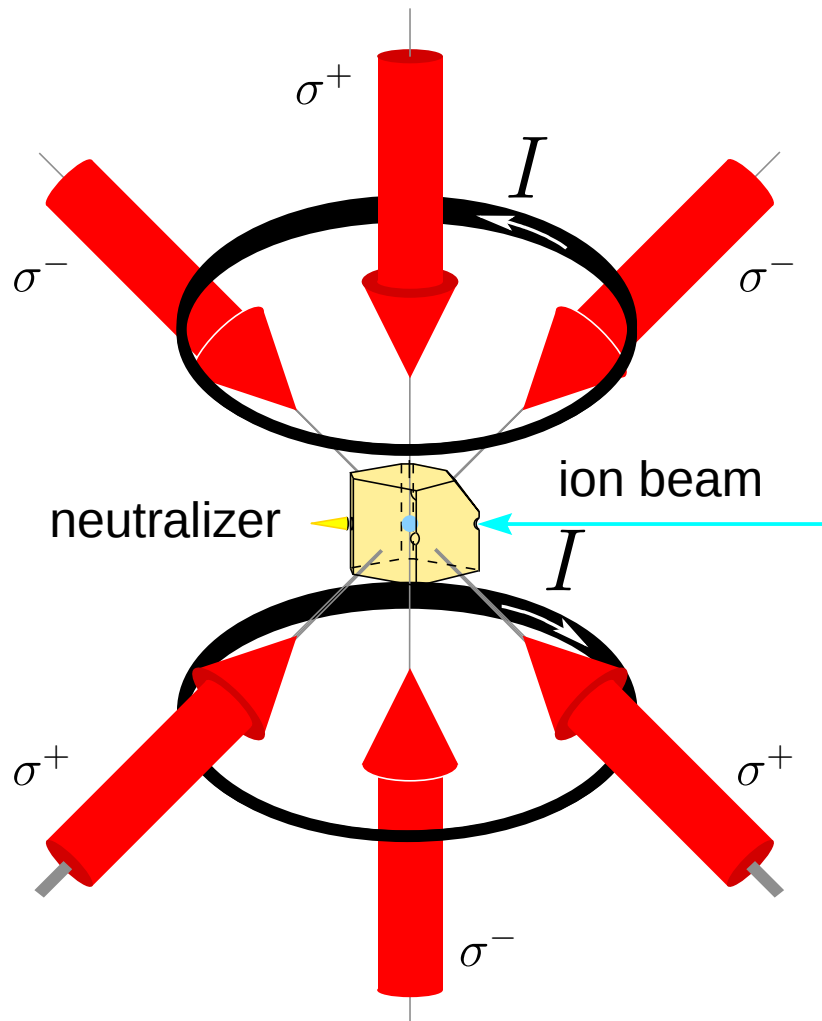
|    |    |    |    |    |    |    |    |    |    |    |    |    |    |    |    |    |    |
|----|----|----|----|----|----|----|----|----|----|----|----|----|----|----|----|----|----|
| H  |    |    |    |    |    |    |    |    |    |    |    |    |    |    |    |    | He |
| Li | Be |    |    |    |    |    |    |    |    |    |    | B  | C  | N  | O  | F  | Ne |
| Na | Mg |    |    |    |    |    |    |    |    |    |    | Al | Si | P  | S  | Cl | Ar |
| K  | Ca | Sc | Ti | V  | Cr | Mn | Fe | Co | Ni | Cu | Zn | Ga | Ge | As | Se | Br | Kr |
| Rb | Sr | Y  | Zr | Nb | Mo | Tc | Ru | Rh | Pd | Ag | Cd | In | Sn | Sb | Te | I  | Xe |
| Cs | Ba | Lu | Hf | Ta | W  | Re | Os | Ir | Pt | Au | Hg | Tl | Pb | Bi | Po | At | Rn |
| Fr | Ra | Lr | Rf | Db | Sg | Bh | Hs | Mt | Ds | Rg |    |    |    |    |    |    |    |

|    |    |    |    |    |    |    |    |    |    |    |    |    |    |
|----|----|----|----|----|----|----|----|----|----|----|----|----|----|
| La | Ce | Pr | Nd | Pm | Sm | Eu | Gd | Tb | Dy | Ho | Er | Tm | Yb |
| Ac | Th | Pa | U  | Np | Pu | Am | Cm | Bk | Cf | Es | Fm | Hd | No |



and *still* the trap is **shallow**...

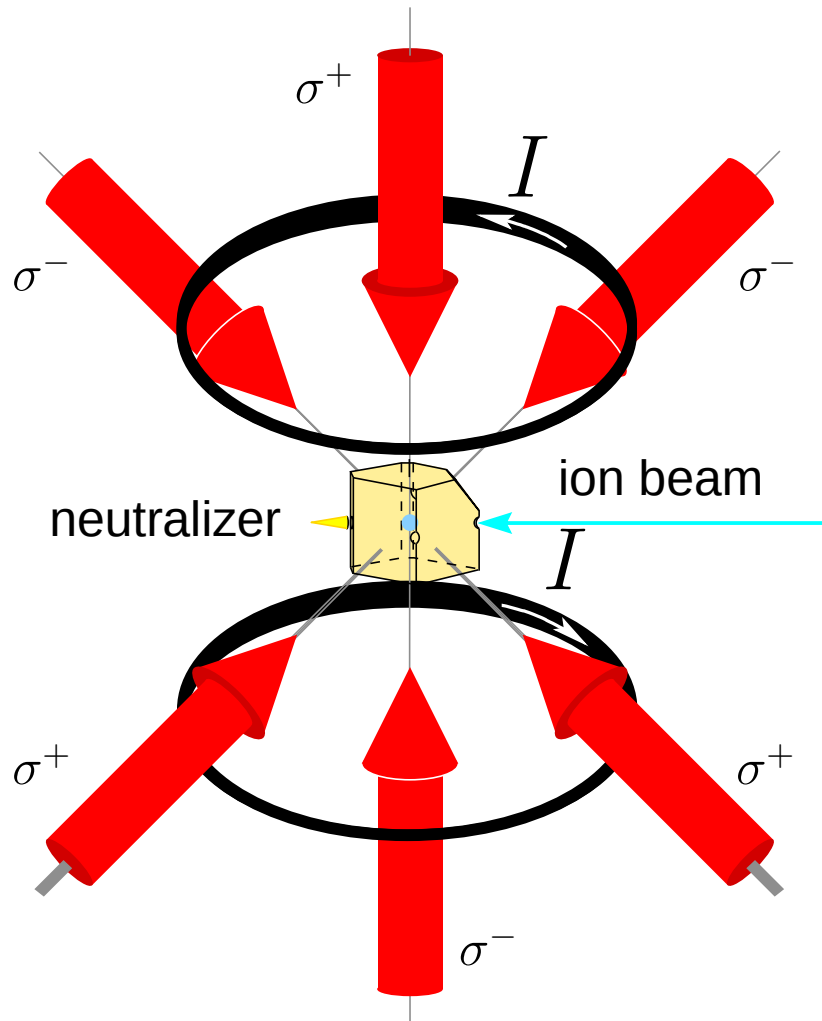
# Still, gotta love it!!



- isomerically selective!

Raab PRL **59** (1987)

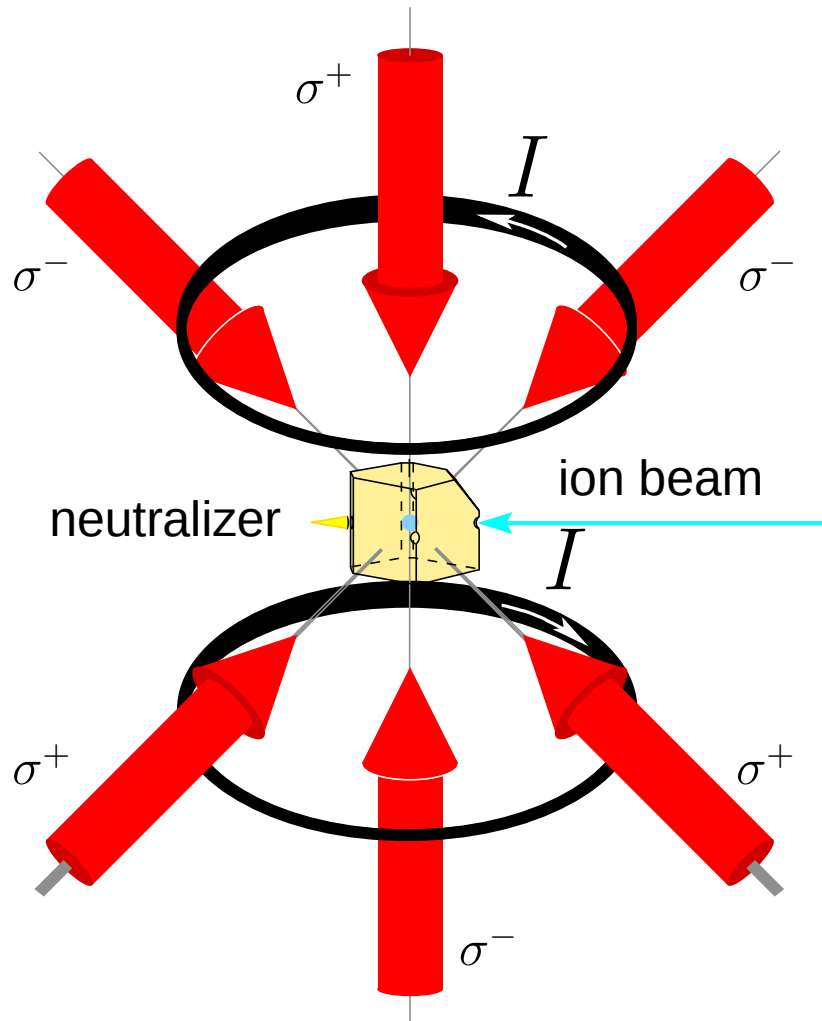
# Still, gotta love it!!



- isomerically selective!
- point-like source!  
( $\lesssim 1 \text{ mm}^3$  FWHM)

Raab PRL **59** (1987)

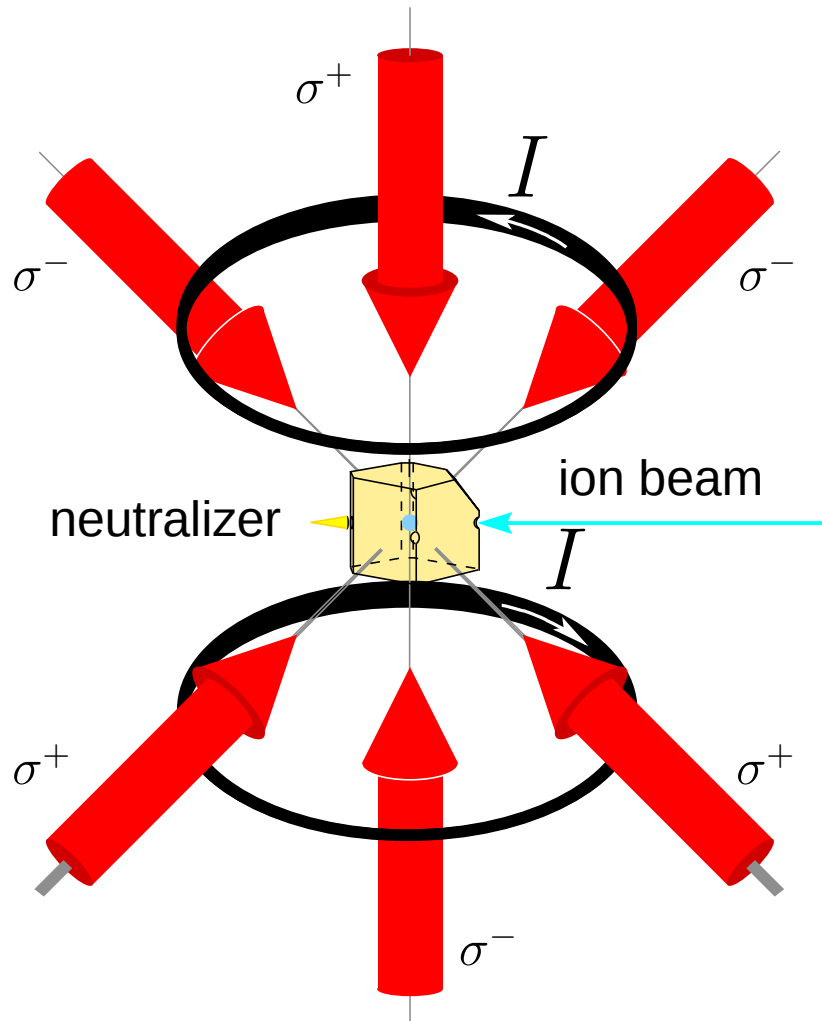
# Still, gotta love it!!



- isomerically selective!
- point-like source!  
( $\lesssim 1 \text{ mm}^3$  FWHM)
- cold atoms! ( $\lesssim 1 \text{ mK}$ )

Raab PRL **59** (1987)

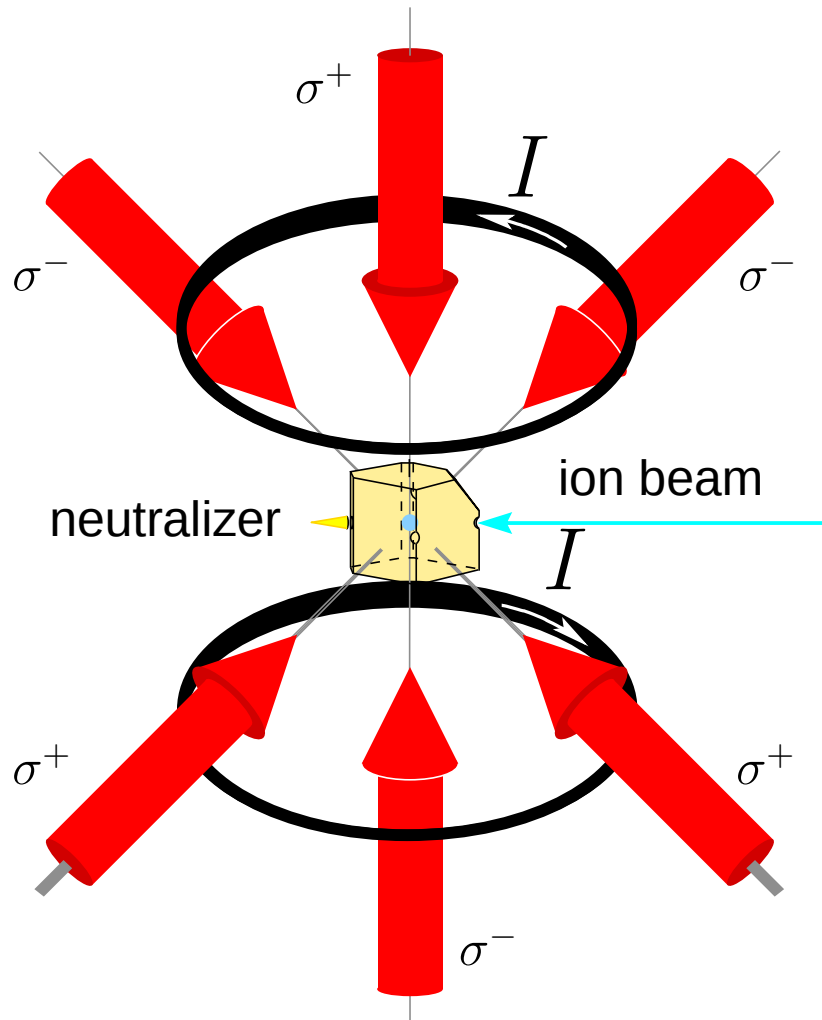
# Still, gotta love it!!



Raab PRL **59** (1987)

- isomerically selective!
- point-like source!  
( $\lesssim 1 \text{ mm}^3$  FWHM)
- cold atoms! ( $\lesssim 1 \text{ mK}$ )
- backing-free source!

# Still, gotta love it!!

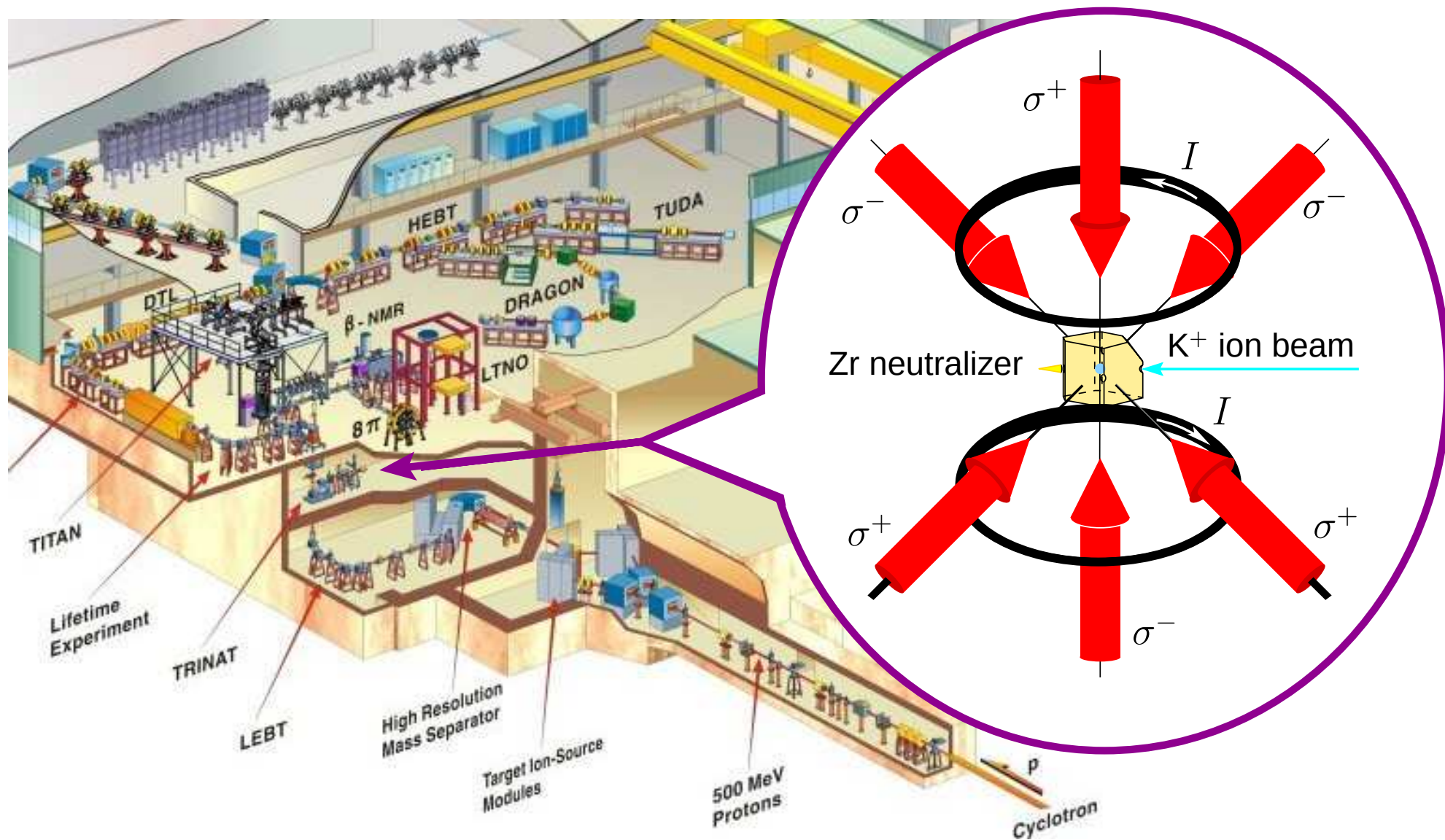


- isomerically selective!
- point-like source!  
( $\lesssim 1 \text{ mm}^3$  FWHM)
- cold atoms! ( $\lesssim 1 \text{ mK}$ )
- backing-free source!

Raab PRL **59** (1987)

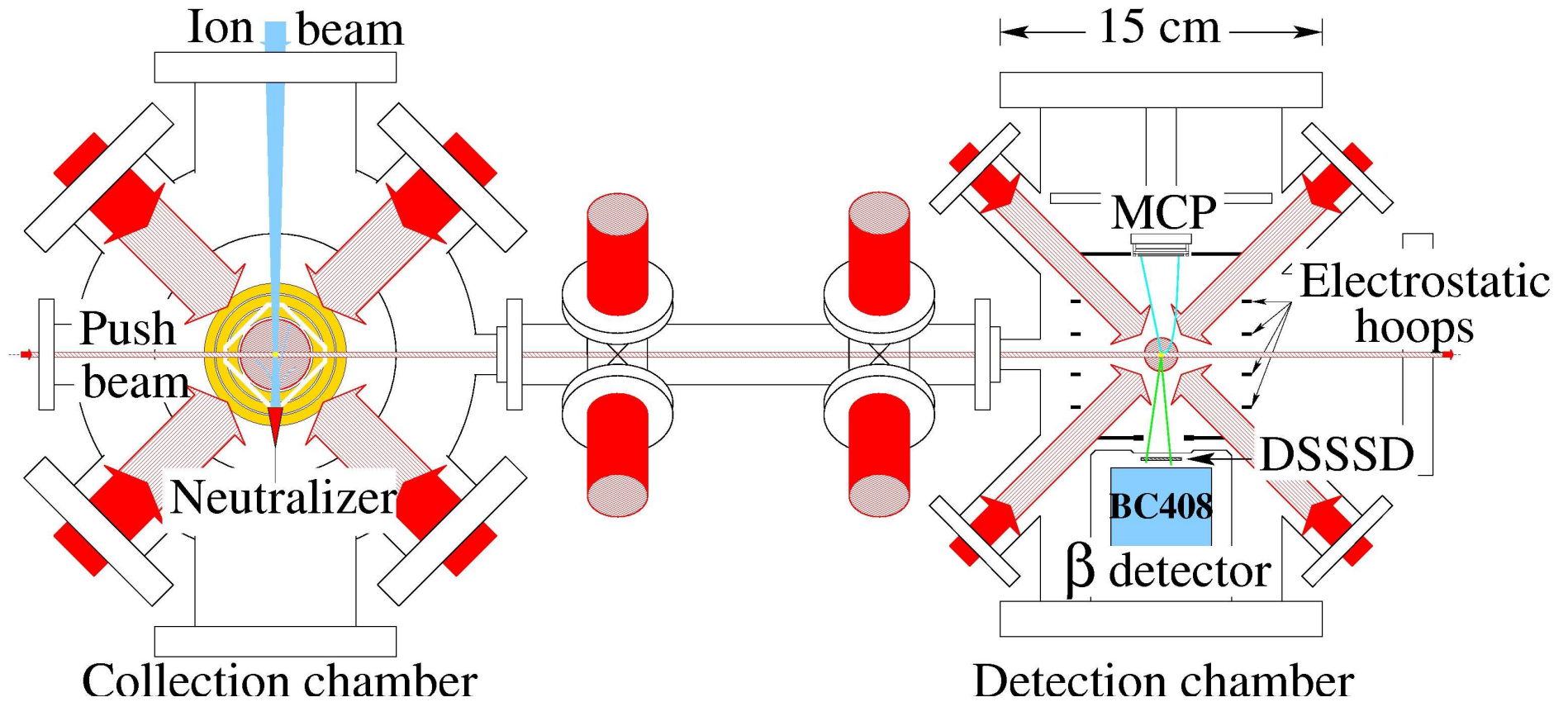
an **ideal** source of radioactives for  $\beta$ -decay experiments!

# Coupling a MOT to ISAC-I

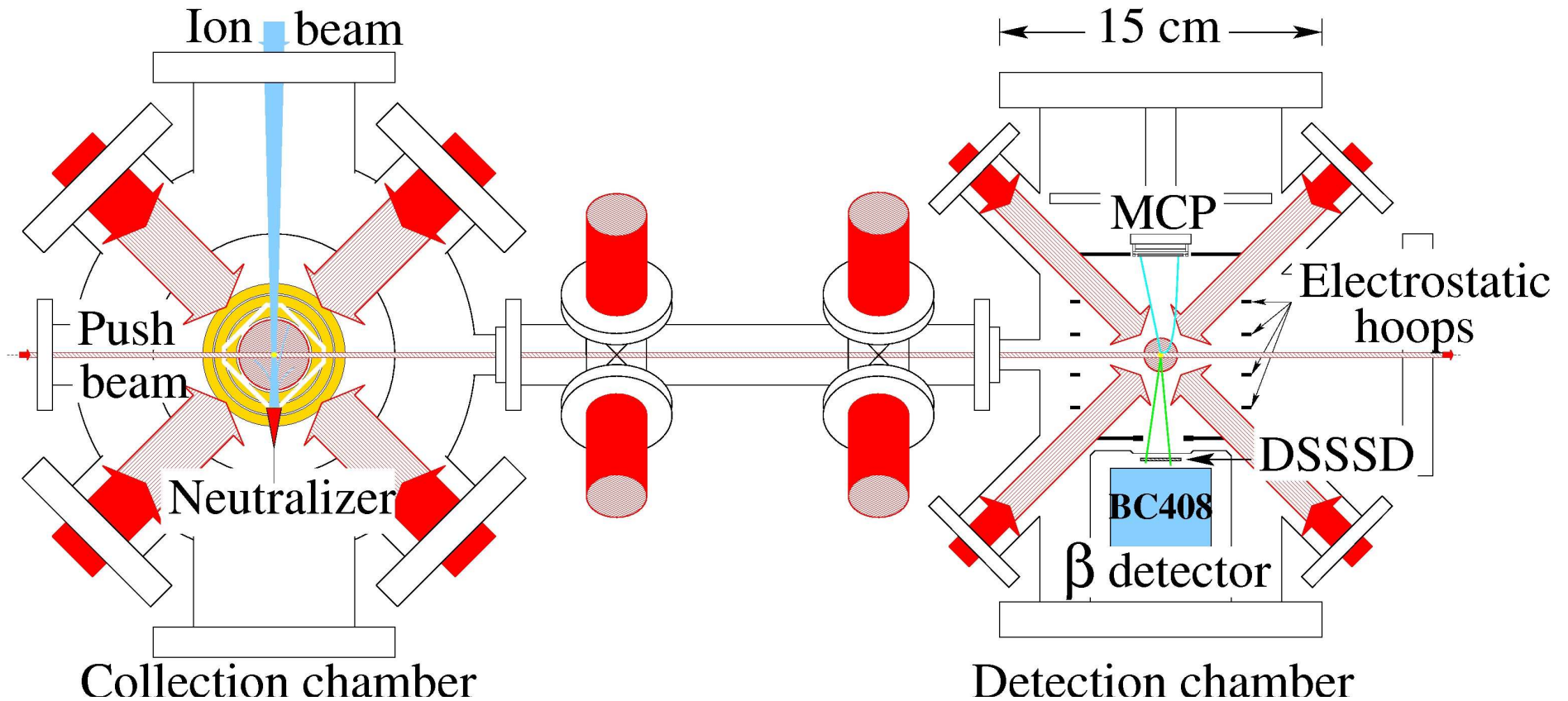


$^{37}\text{K}$  yield with 40  $\mu\text{A}$  on TiC #1:  $6 \times 10^7/\text{s}$

# Double-MOT system



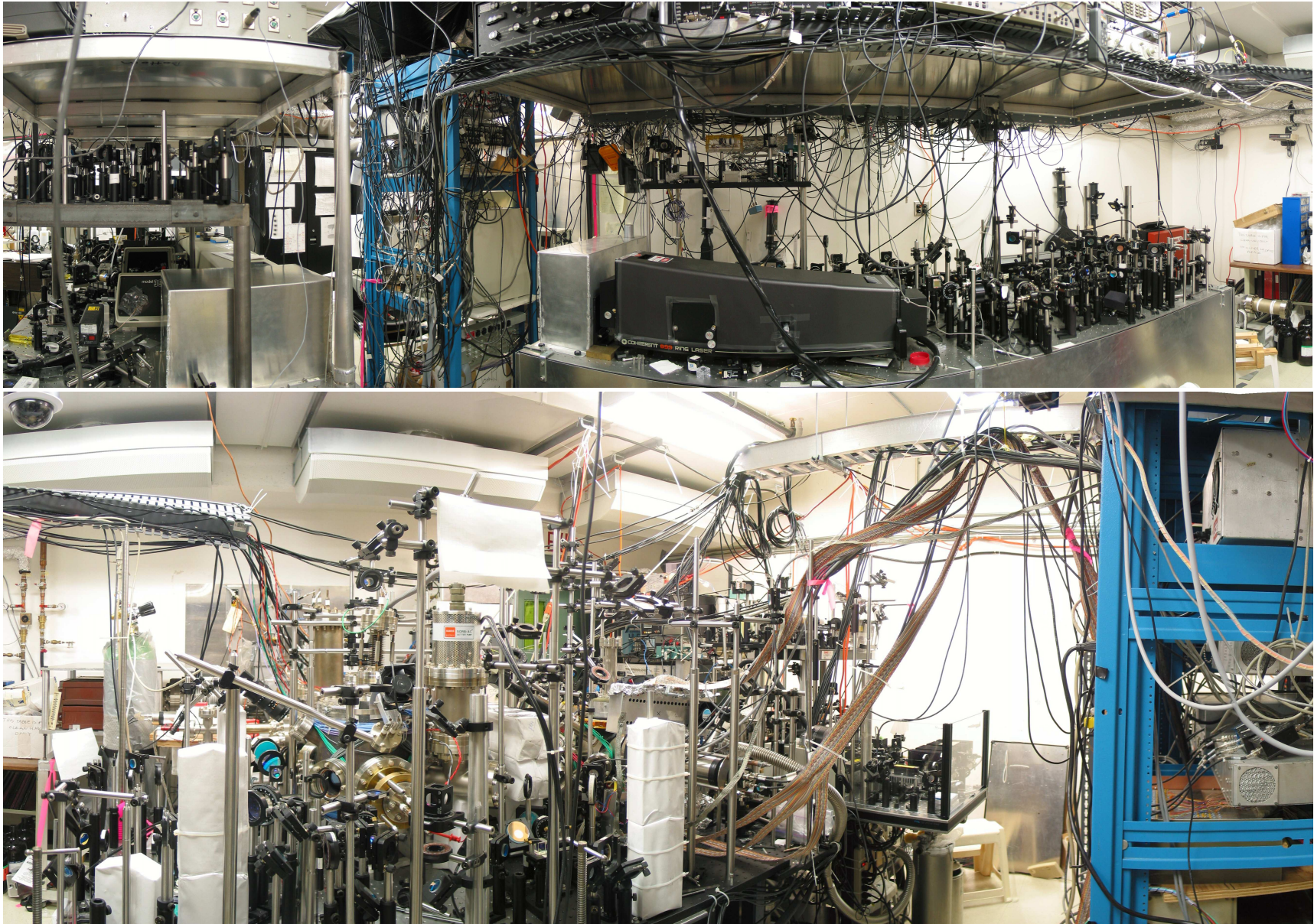
# Double-MOT system



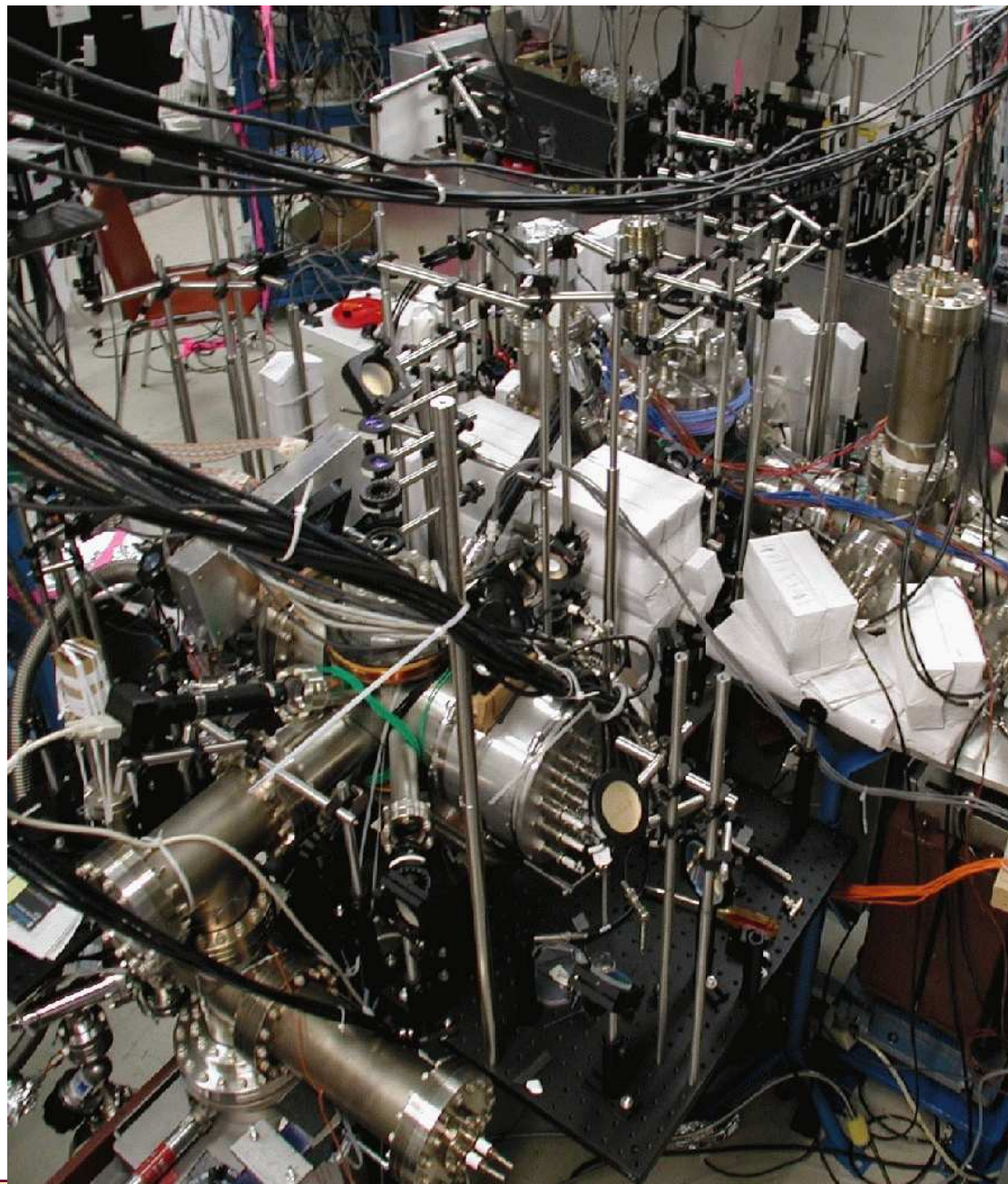
Traps provide a backing-free, cold ( $\lesssim 1$  mK), localized ( $\lesssim 1$  mm<sup>3</sup>) source of short-lived radioactive atoms

Detect  $p_\beta$  and  $p_{\text{recoil}} \Rightarrow$  deduce  $p_\nu$ !

# *The TRINAT lab*



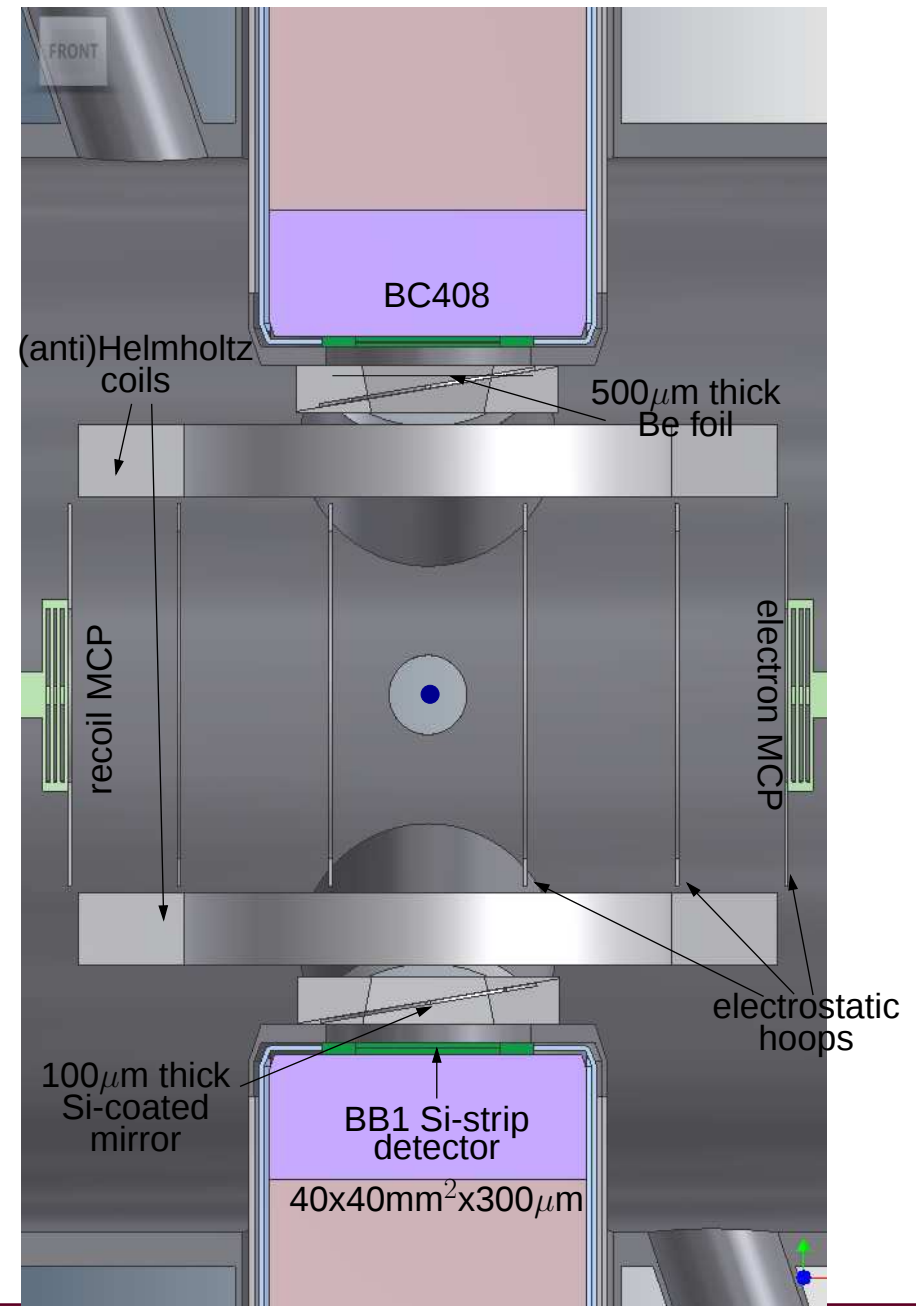
# *The TRINAT lab*



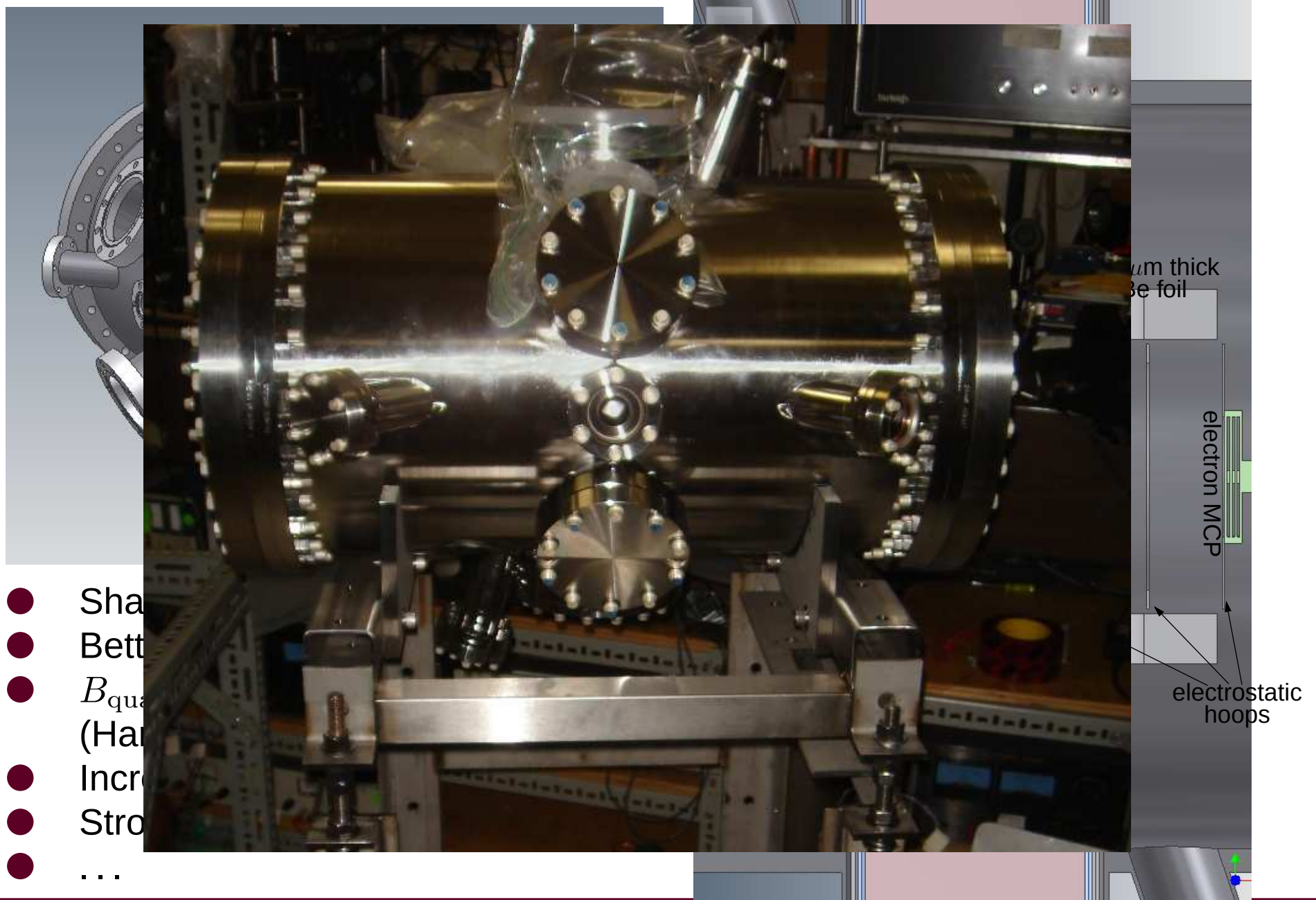
# The new chamber



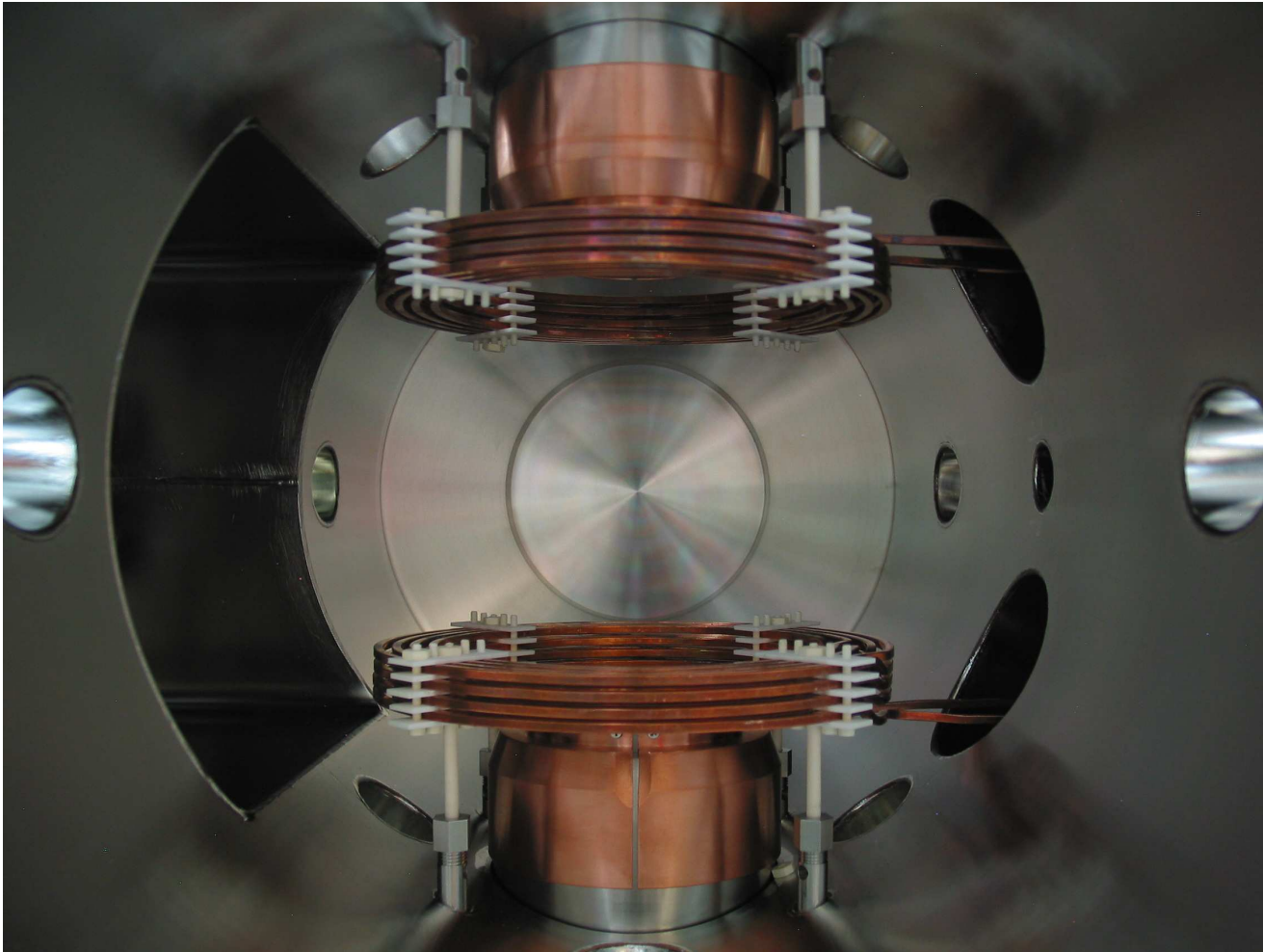
- Shake-off  $e^-$  detection
- Better control of OP beams
- $B_{\text{quad}} \rightarrow B_{\text{OP}}$  quickly: AC-MOT (Harvery & Murray, PRL **101** (2008))
- Increased  $\beta$ /recoil solid angles
- Stronger  $E$ -field
- ...



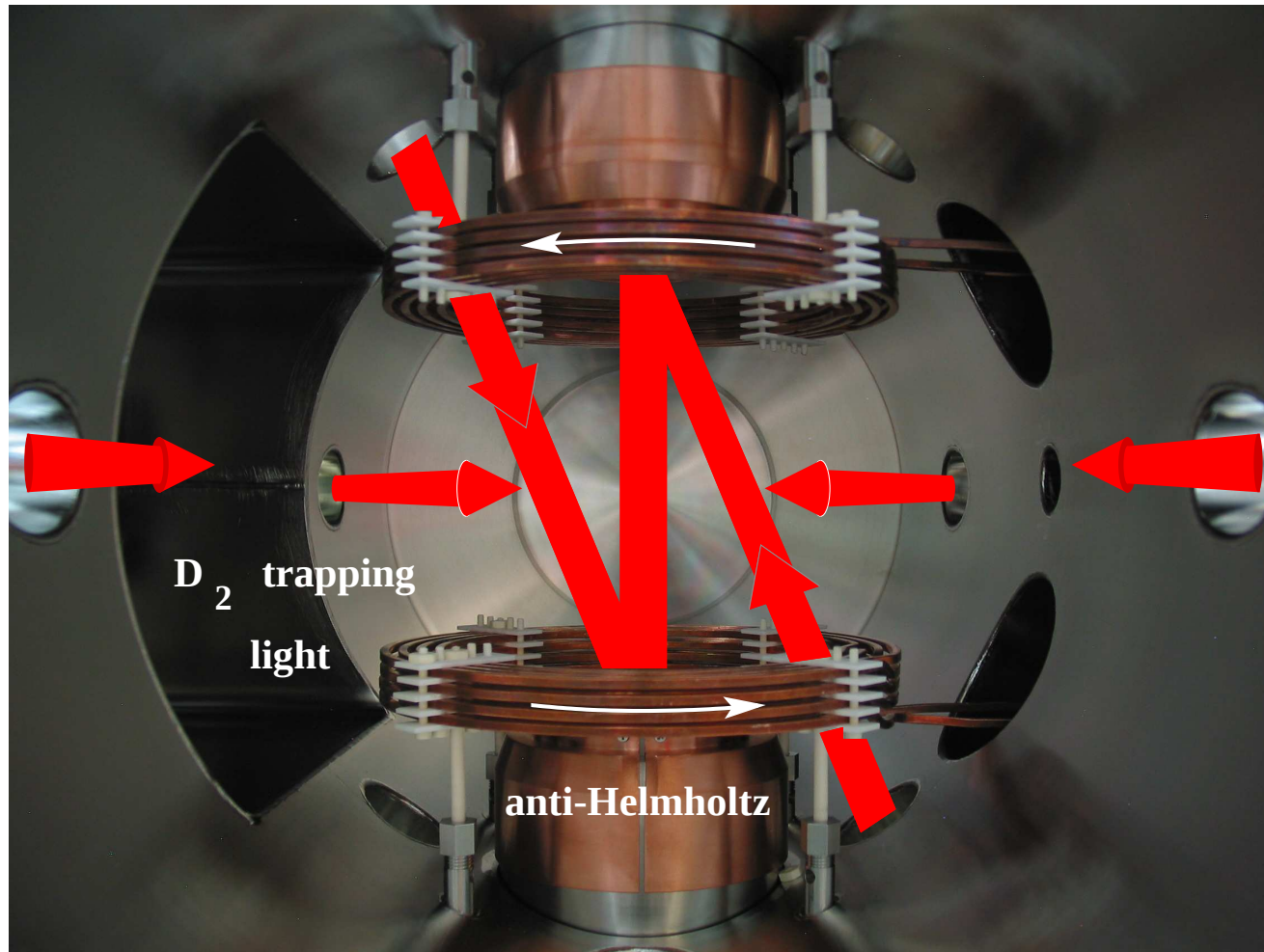
# The new chamber



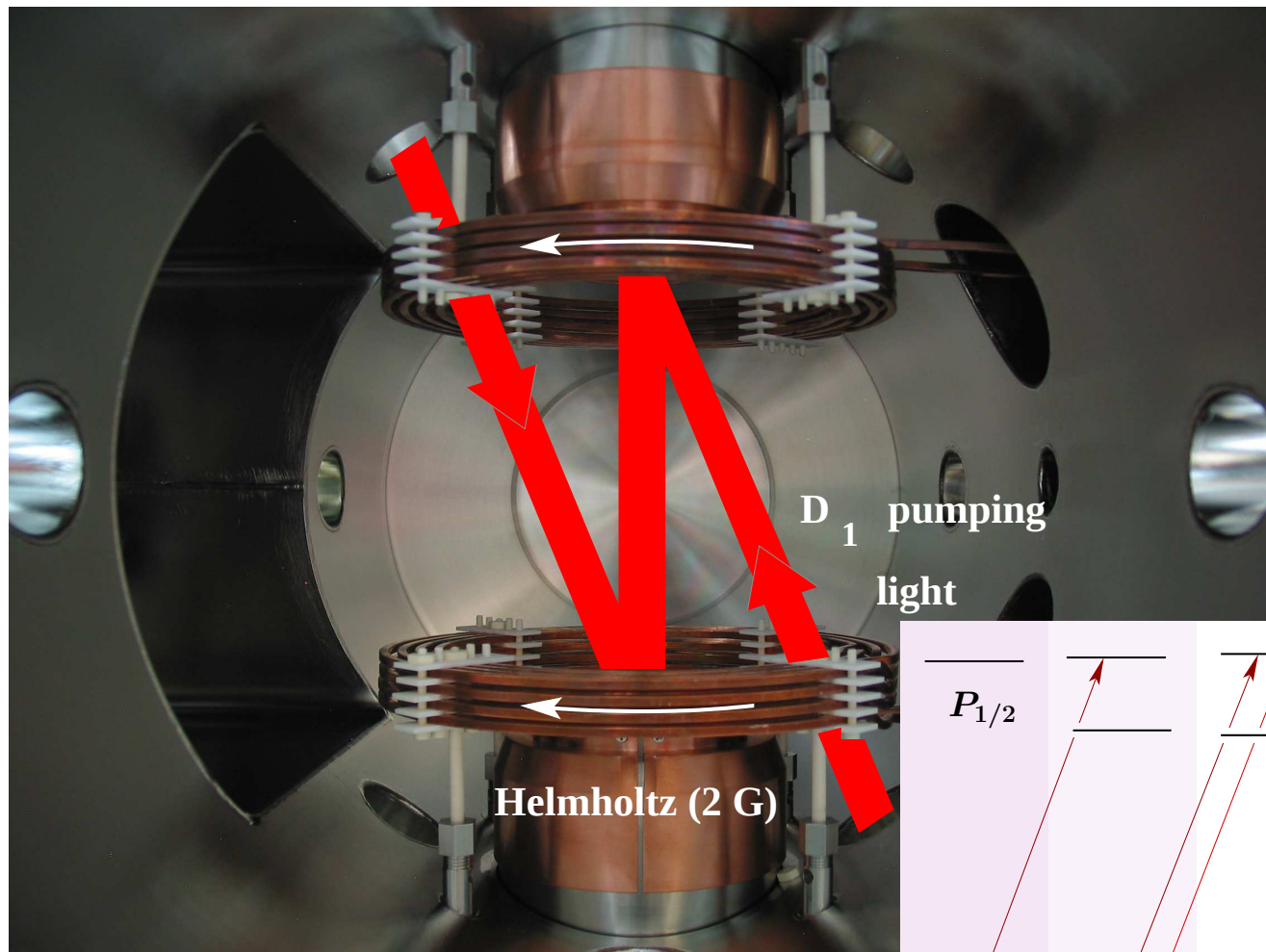
# *Outline of polarized experiment*



# Outline of polarized experiment



# Outline of polarized experiment



$$\mathbf{F} = \mathbf{I} + \mathbf{J}$$

$$I = \frac{3}{2}$$

$$J = \frac{1}{2}$$

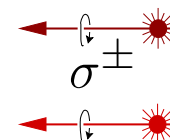
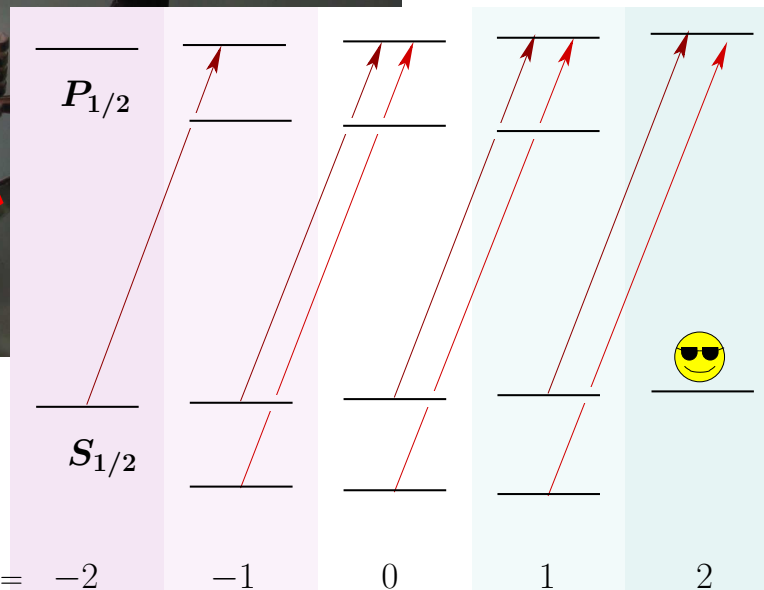
$$m_F = -2$$

$$-1$$

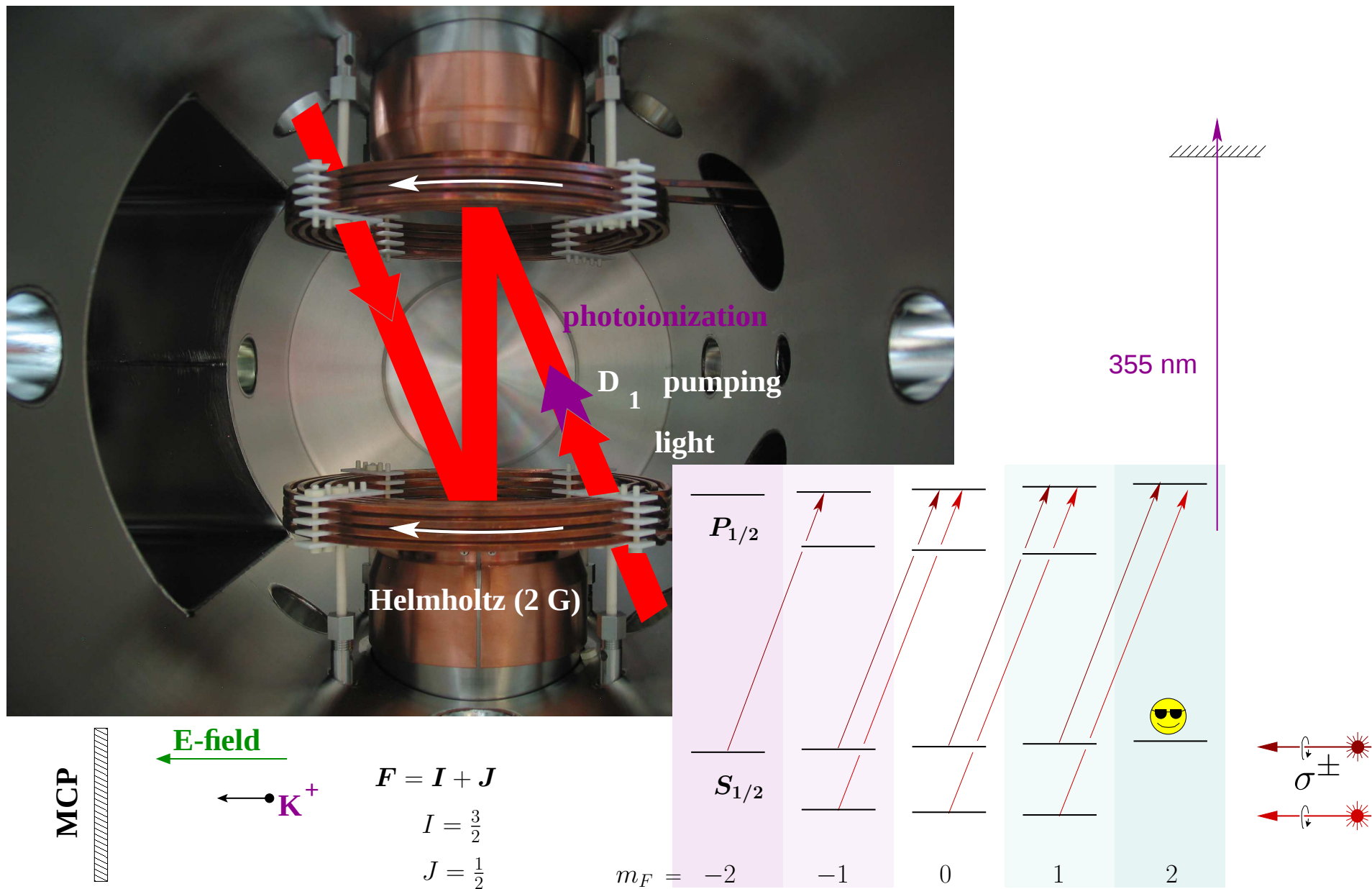
$$0$$

$$1$$

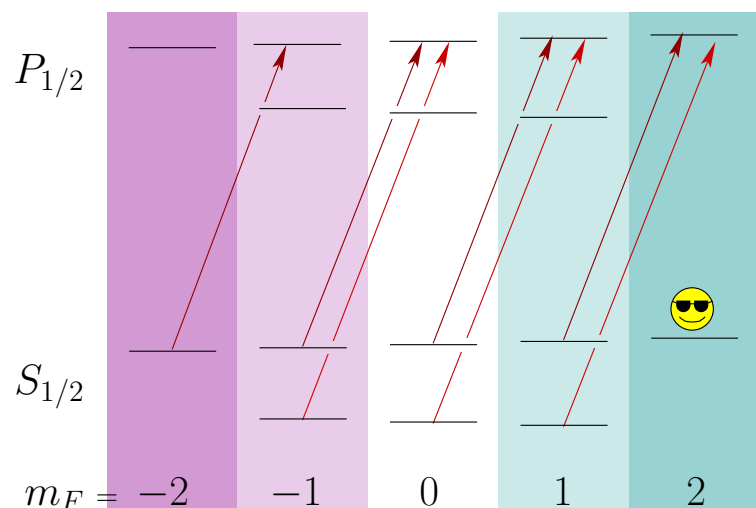
$$2$$



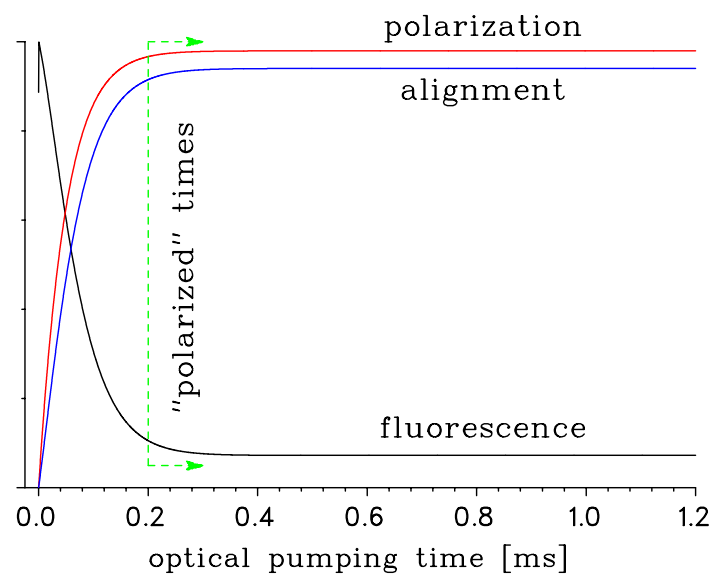
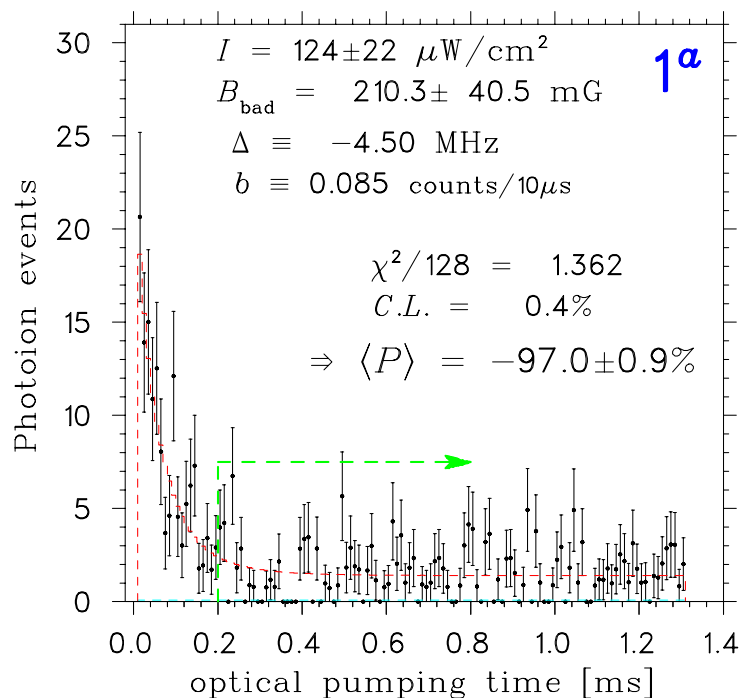
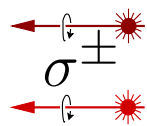
# Outline of polarized experiment



# Atomic measurement of $P$



deduce  $P$  based on a model of the excited state populations:



$$\Rightarrow P_{\text{nucl}} = 96.74 \pm 0.53^{+0.19}_{-0.73}$$

# Atomic measurement of $P$

$P_{1/2}$

PRL **101**, 173201 (2008)

PHYSICAL REVIEW LETTERS

week ending  
24 OCTOBER 2008

for MOT currents rapidly switched to zero, the induced eddy currents continue to produce  $\mathbf{B}$  fields until they too reduce to zero.

In practice the  $\mathbf{B}$  field due to the MOT takes  $\sim 10$  ms to reduce to  $< 10^{-7}$  T, this time depending on the proximity of conductors to the coils, their shape, and resistivity. During this time, a large fraction of trapped atoms escape, resulting in a cold atom density that rapidly falls to zero. Losses can be reduced by leaving the cooling lasers on to create an optical molasses (if this does not interfere with the experiment); however, the loss problems remain. The comparatively long time taken for the  $\mathbf{B}$  field to decay also reduces data accumulation rates, since the repetition rate is then only  $\sim 50$  Hz.

It is clearly advantageous to eliminate these constraints. Several methods have been attempted, including shaping the dc MOT driving current at switchoff to try to cancel fields due to eddy currents [10]. This technique is complicated and requires different currents when spectrometer

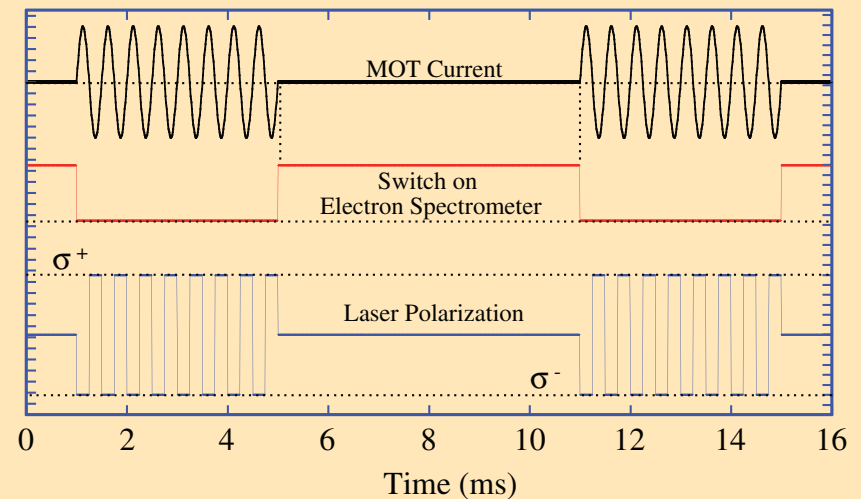
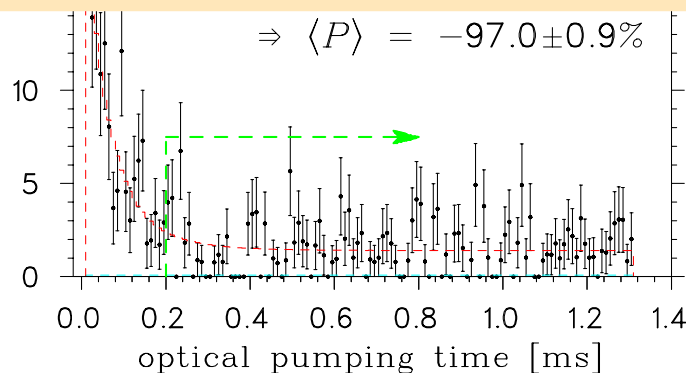


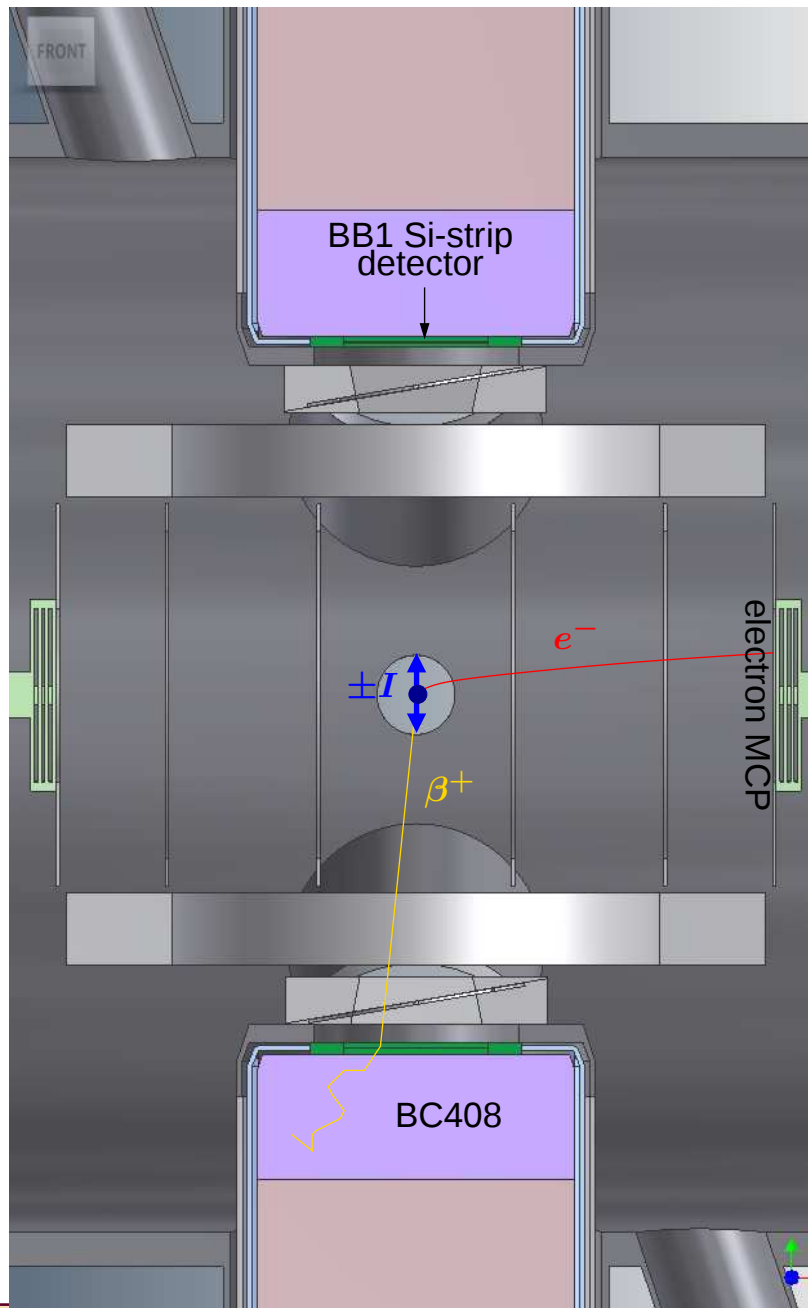
FIG. 1 (color online). The switching configuration for the ac MOT. The MOT is driven by an alternating supply, so that the net induced current in conductors surrounding the MOT coils is zero. The polarization of the six trapping laser beams is switched at the same rate as the MOT current, so as to maintain trapping. Experiments using charged particles are conducted during the time the MOT current is zero.

Photoion events



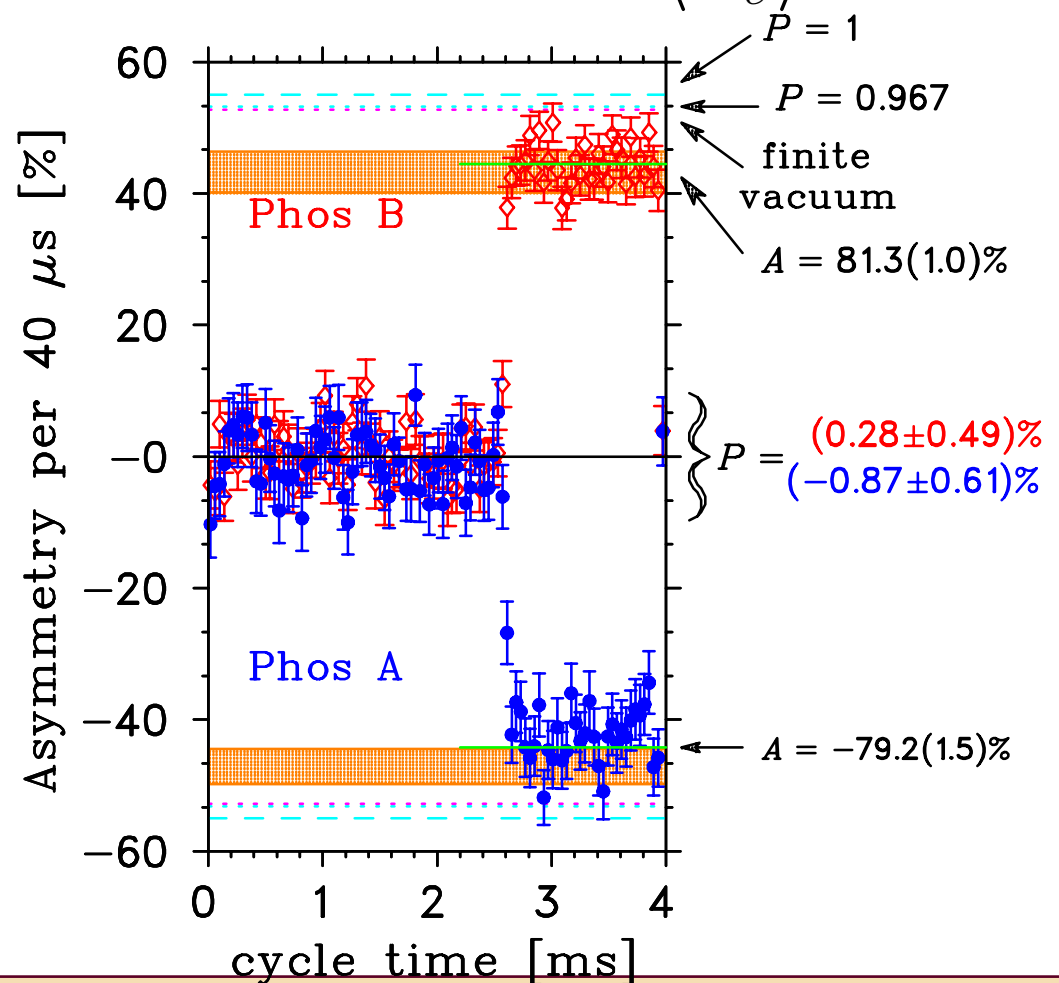
$$\Rightarrow P_{\text{nucl}} = 96.74 \pm 0.53^{+0.19}_{-0.73}$$

# Measuring $A_\beta$ — Fall 2012(?)



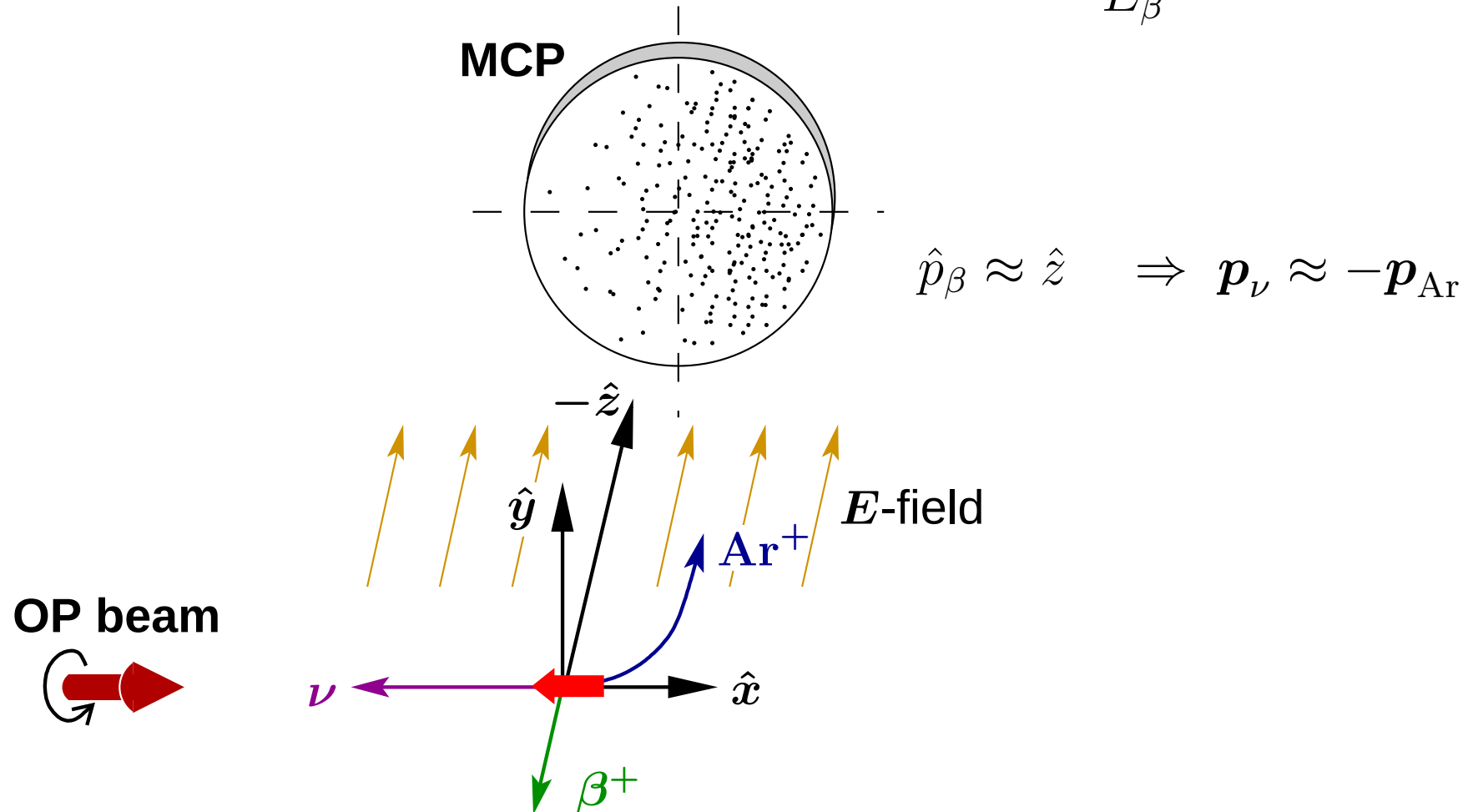
$$\text{Asymmetry} = \frac{N(\sigma^+) - N(\sigma^-)}{N(\sigma^+) + N(\sigma^-)}$$

$$\sim P A_\beta \left\langle \frac{p_e}{E_e} \right\rangle$$



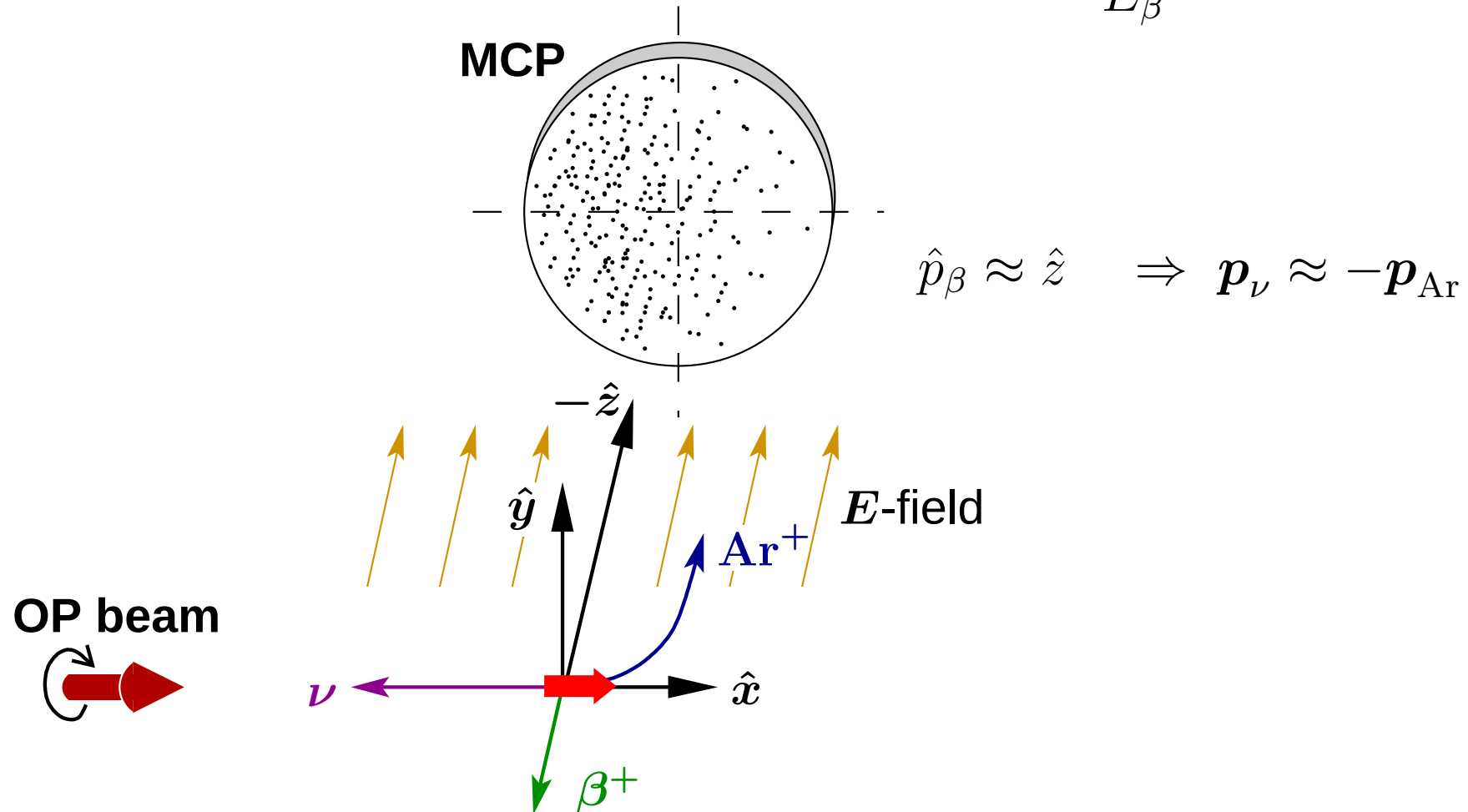
# Measuring $B_\nu$ (and $D$ )

$$dW \sim \textcolor{blue}{P}\textcolor{green}{B}_\nu \hat{p}_\nu \cdot \hat{i} + \textcolor{blue}{P}\textcolor{red}{D} \frac{\hat{i} \cdot (\mathbf{p}_\beta \times \hat{p}_\nu)}{E_\beta}$$



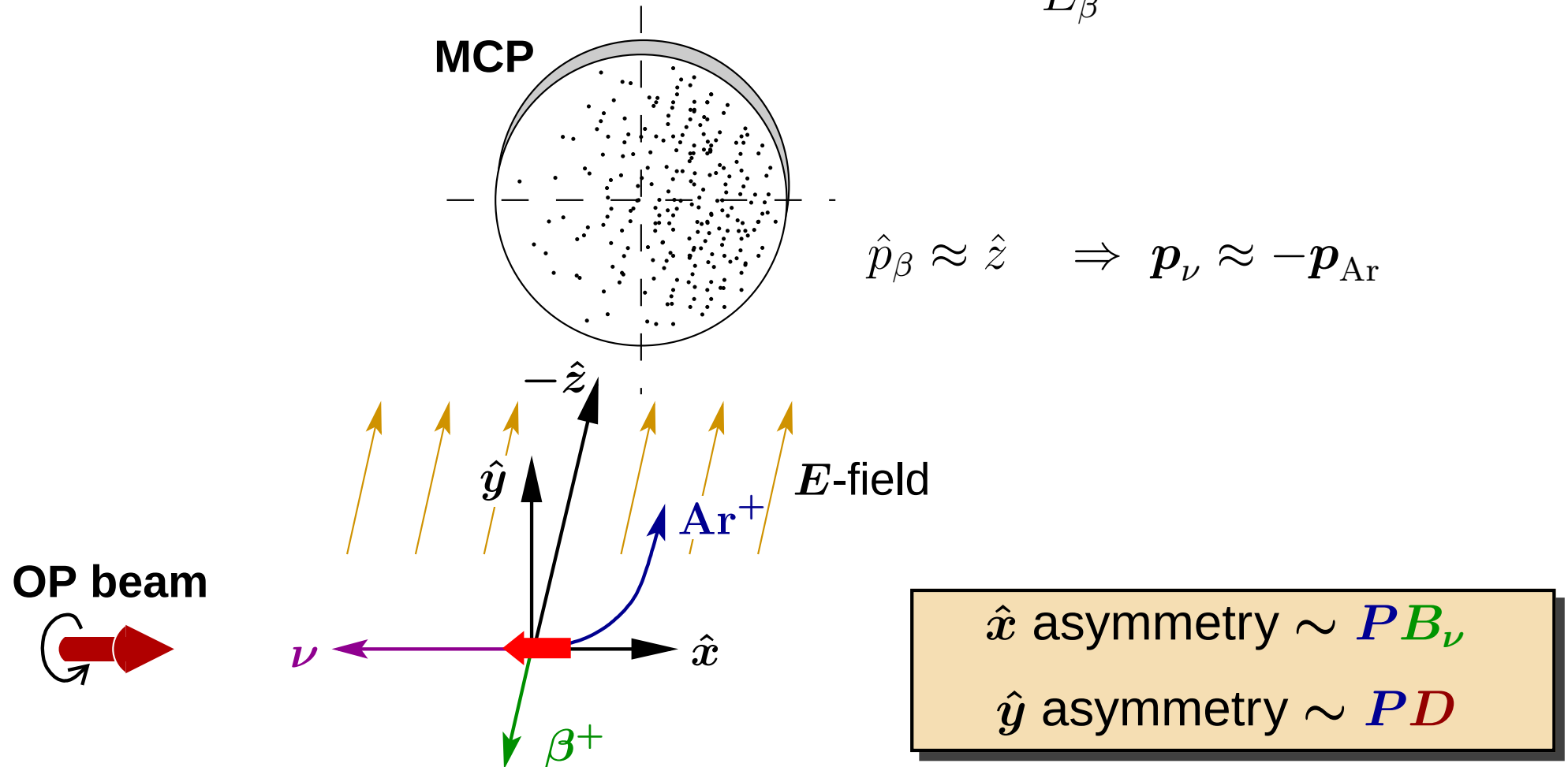
# Measuring $B_\nu$ (and $D$ )

$$dW \sim \textcolor{blue}{P}\textcolor{green}{B}_\nu \hat{p}_\nu \cdot \hat{i} + \textcolor{blue}{P}\textcolor{red}{D} \frac{\hat{i} \cdot (\mathbf{p}_\beta \times \hat{p}_\nu)}{E_\beta}$$

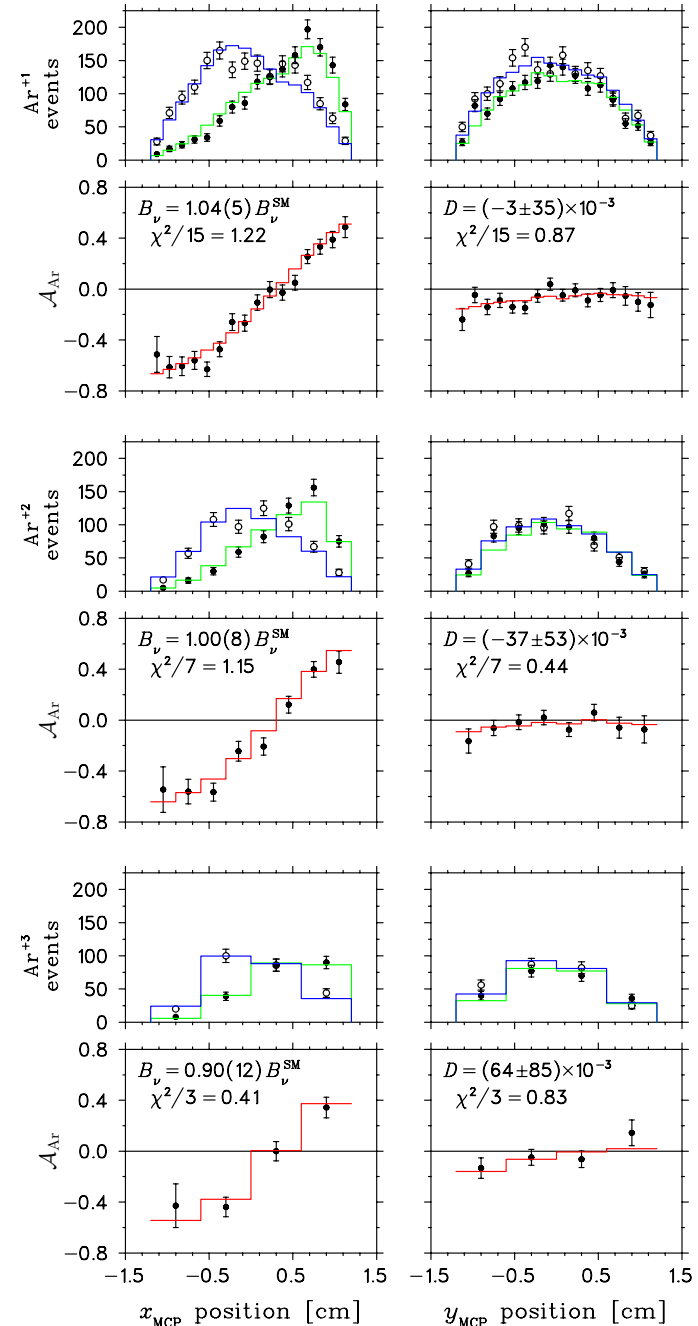
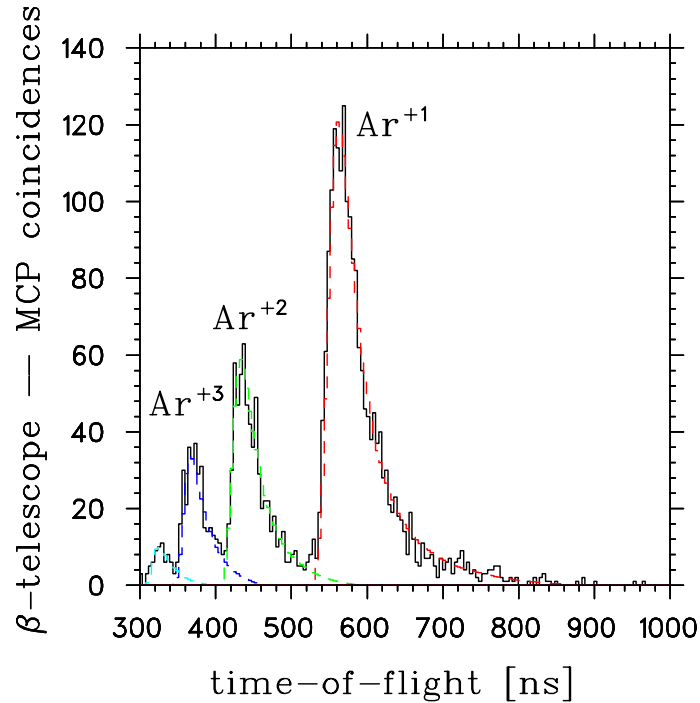


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# The neutrino asymmetry measurement



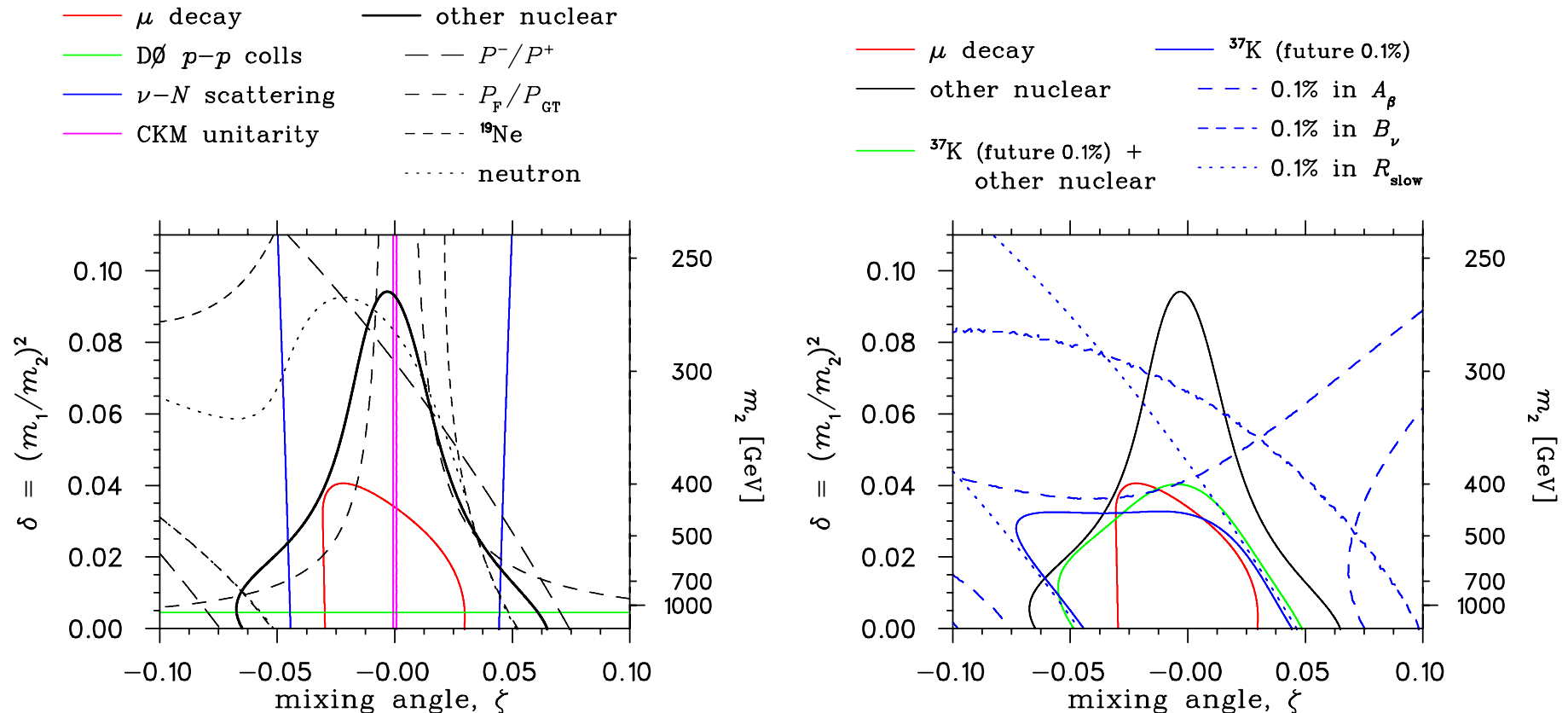
$$1^{\text{st}} : \langle B_\nu \rangle = (0.995 \pm 0.040)_{(\text{stat})} B_\nu^{\text{SM}}$$

$$2^{\text{nd}} : \langle B_\nu \rangle = (0.975 \pm 0.031)_{(\text{stat})} B_\nu^{\text{SM}}$$

$$\Rightarrow \boxed{B_\nu = 0.981(26)(17) B_\nu^{\text{SM}}}$$

(Melconian, PLB **649** (2007) 370)

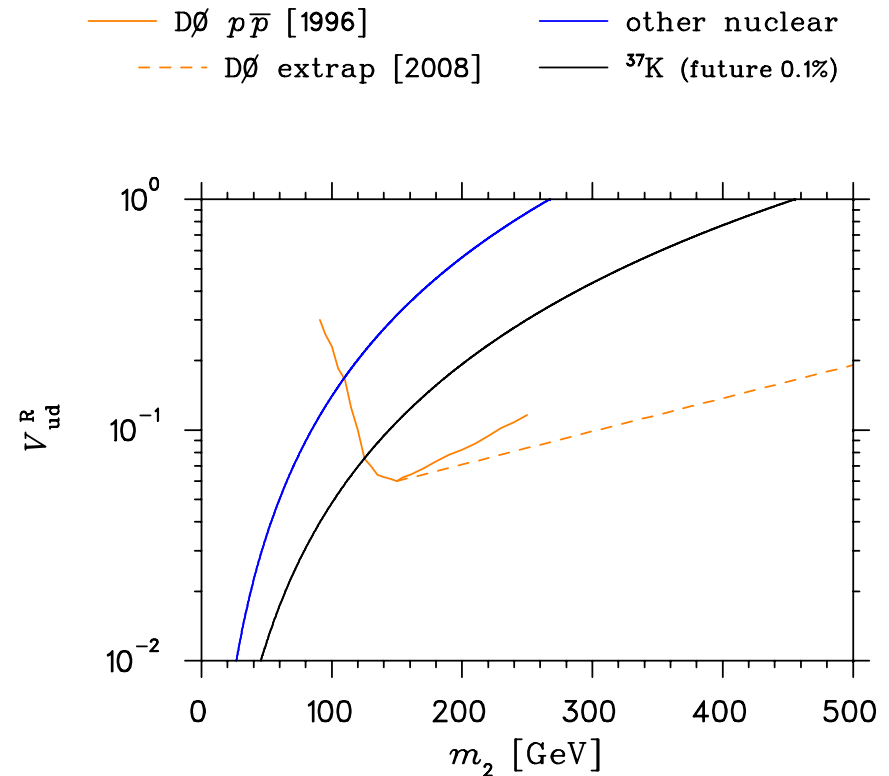
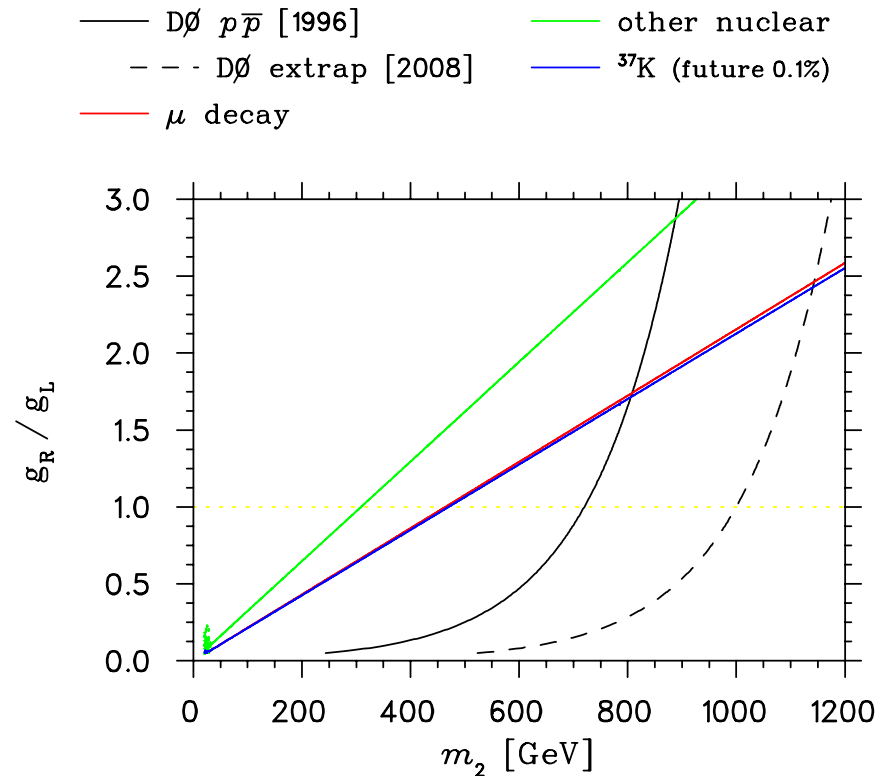
# Goal, in terms of RHCs



**Expected limits if  $A_\beta$ ,  $B_\nu$  and  $R_{\text{slow}}$  all measured to 0.1%**

see Profumo, Ramsey-Musolf and Tulin, PRD **75** (2007) 075017

# Beyond the minimal L-R symmetric model



(adapted from Thomas *et al.*, Nucl Phys A **694**; see also Severijns, Beck and Naviliat-Cuncic, Rev Mod Phys **78** (2006))

different experiments are complementary

# Summary

- SM is fantastic, but **incomplete**
- many **exciting avenues** to find more complete model
- **needed:** precision measurement of correlation parameters
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Don't get married five days before you're supposed to give a talk! ;-)

# *Main collaborators/thanks*

Cyclotron Institute: P. Shidling, **B. Fenker**, **M. Mehlman**, **S. Behling**, J.C. Hardy, R. Tribble, G. Chubaryan, L. Trache, staff, ...

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